

Mesure du chaos

Haris Skokos

Astronomie et Systèmes Dynamiques, IMCCE,
Observatoire de Paris, Paris, France

E-mail: hskokos@imcce.fr

URL: <http://www.imcce.fr/~hskokos/>

Outline

- Basic concepts of Hamiltonian systems and symplectic maps
 - ✓ Poincaré Surface of Section
 - ✓ Variational equations
 - ✓ Lyapunov exponents
 - ✓ Chaos detection techniques
- Smaller ALignment Index – SALI
 - ✓ Definition
 - ✓ Behavior for chaotic and regular motion
 - ✓ Applications
- Generalized ALignment Index – GALI
 - ✓ Definition - Relation to SALI
 - ✓ Behavior for chaotic and regular motion
 - ✓ Applications
 - ✓ Global dynamics

Autonomous Hamiltonian systems

Consider an **N degree of freedom** autonomous Hamiltonian system having a Hamiltonian function of the form:

$$H(\underbrace{q_1, q_2, \dots, q_N}_{\text{positions}}, \underbrace{p_1, p_2, \dots, p_N}_{\text{momenta}})$$

The time evolution of an orbit (trajectory) with initial condition

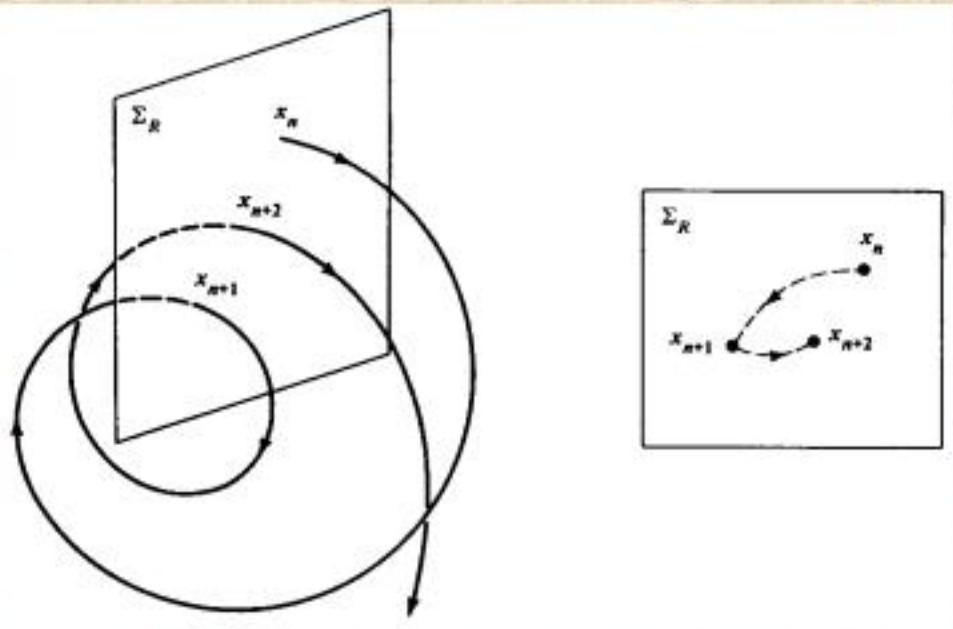
$$P(0) = (q_1(0), q_2(0), \dots, q_N(0), p_1(0), p_2(0), \dots, p_N(0))$$

is governed by the **Hamilton's equations of motion**

$$\frac{dp_i}{dt} = -\frac{\partial H}{\partial q_i}, \quad \frac{dq_i}{dt} = \frac{\partial H}{\partial p_i}$$

Poincaré Surface of Section (PSS)

We can constrain the study of an $N+1$ degree of freedom Hamiltonian system to a **2N-dimensional subspace** of the general phase space.



Lieberman & Lichtenberg, 1992, *Regular and Chaotic Dynamics*, Springer.

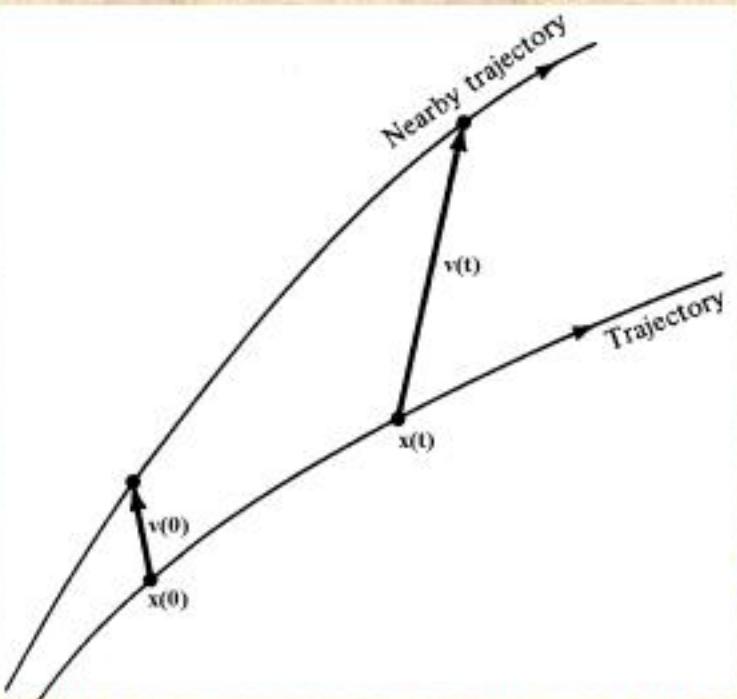
In general we can assume a PSS of the form $q_{N+1} = \text{constant}$. Then only variables $q_1, q_2, \dots, q_N, p_1, p_2, \dots, p_N$ are needed to describe the evolution of an orbit on the PSS, since p_{N+1} can be found from the Hamiltonian.

In this sense an $N+1$ degree of freedom Hamiltonian system corresponds to a **2N-dimensional symplectic map**.

Variational Equations

We use the notation $x = (q_1, q_2, \dots, q_N, p_1, p_2, \dots, p_N)^T$. The **deviation vector** from a given orbit is denoted by

$$v = (dx_1, dx_2, \dots, dx_n)^T, \text{ with } n=2N$$



The time evolution of v is given by the so-called **variational equations**:

$$\frac{dv}{dt} = -J \cdot P \cdot v$$

where

$$J = \begin{pmatrix} 0_N & -I_N \\ I_N & 0_N \end{pmatrix}, \quad P_{ij} = \frac{\partial^2 H}{\partial x_i \partial x_j} \quad i, j = 1, 2, \dots, n$$

Benettin & Galgani, 1979, in Laval and Gressillon (eds.), op cit, 93

Example (Hénon-Heiles system)

$$H = \frac{1}{2}(p_x^2 + p_y^2) + \frac{1}{2}(x^2 + y^2) + x^2y - \frac{1}{3}y^3$$

Hamilton's equations of motion:

$$\frac{dp_i}{dt} = -\frac{\partial H}{\partial q_i}, \quad \frac{dq_i}{dt} = \frac{\partial H}{\partial p_i} \Rightarrow \begin{cases} \dot{x} = p_x \\ \dot{y} = p_y \\ \dot{p}_x = -x - 2xy \\ \dot{p}_y = -y - x^2 + y^2 \end{cases}$$

In order to get the variational equations we **linearize** the above equations by substituting x, y, p_x, p_y with $x+v_1, y+v_2, p_x+v_3, p_y+v_4$ where $v=(v_1, v_2, v_3, v_4)$ is the deviation vector. So we get:

$$\dot{p}_x + \dot{v}_3 = -x - v_1 - 2(x + v_1)(y + v_2) \Rightarrow$$

~~$$\dot{p}_x + \dot{v}_3 = -x - v_1 - 2xy - 2xv_2 - 2yv_1 - 2vv_2 \Rightarrow$$~~

$$\dot{v}_3 = -v_1 - 2yv_1 - 2xv_2$$

Example (Hénon-Heiles system)

Variational equations:

$$\frac{dv}{dt} = -J \cdot P \cdot v$$

$$\begin{pmatrix} \dot{v}_1 \\ \dot{v}_2 \\ \dot{v}_3 \\ \dot{v}_4 \end{pmatrix} = - \begin{pmatrix} 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \\ 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{pmatrix} \begin{pmatrix} 1+2y & 2x & 0 & 0 \\ 2x & 1-2y & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} v_1 \\ v_2 \\ v_3 \\ v_4 \end{pmatrix}$$

$$\dot{v}_1 =$$

$$v_3$$

$$\dot{x} = p_x$$

$$\dot{v}_2 =$$

$$v_4$$

$$\dot{y} = p_y$$

$$\boxed{\dot{v}_3 = -v_1 - 2xv_2 - 2yv_1}$$



$$\dot{p}_x = -x - 2xy$$

$$\dot{v}_4 = -v_2 - 2xv_1 + 2yv_2$$

$$\dot{p}_y = -y - x^2 + y^2$$

Complete set of equations

Symplectic Maps

Consider an **n-dimensional symplectic map T**. In this case we have **discrete time**.

The evolution of an **orbit** with initial condition

$$P(0) = (x_1(0), x_2(0), \dots, x_n(0))$$

is governed by the **equations of map T**

$$P(i+1) = T P(i) , i=0,1,2,\dots$$

The evolution of an initial **deviation vector**

$$v(0) = (dx_1(0), dx_2(0), \dots, dx_n(0))$$

is given by the corresponding **tangent map**

$$v(i+1) = \left. \frac{\partial T}{\partial P} \right|_i \cdot v(i) , i=0,1,2,\dots$$

Example – 2D map

Equations of the map:

$$\begin{pmatrix} \mathbf{x}'_1 \\ \mathbf{x}'_2 \end{pmatrix} = \mathbf{T} \begin{pmatrix} \mathbf{x}_1 \\ \mathbf{x}_2 \end{pmatrix} \Rightarrow \begin{aligned} \mathbf{x}'_1 &= \mathbf{x}_1 + \mathbf{x}_2 \\ \mathbf{x}'_2 &= \mathbf{x}_2 - v \sin(\mathbf{x}_1 + \mathbf{x}_2) \end{aligned} \quad (\text{mod } 2\pi)$$

Tangent map:

$$v(i+1) = \frac{\partial \mathbf{T}}{\partial \mathbf{P}} \Big|_i \cdot v(i)$$
$$\begin{pmatrix} \mathbf{dx}'_1 \\ \mathbf{dx}'_2 \end{pmatrix} = \begin{pmatrix} 1 & 1 \\ -v \cos(\mathbf{x}_1 + \mathbf{x}_2) & 1 - v \cos(\mathbf{x}_1 + \mathbf{x}_2) \end{pmatrix} \begin{pmatrix} \mathbf{dx}_1 \\ \mathbf{dx}_2 \end{pmatrix}$$

Chaos detection methods

- Analysis of the orbit under study
 - ✓ Frequency Analysis
 - ✓ Low frequency power (LFP)
 - ✓ 0-1 test
 - ✓ Patterns method
- Study of the evolution of deviation vectors
 - ✓ Lyapunov exponents
 - ✓ SALI – GALI_k
 - ✓ Fast Lyapunov Indicator (FLI) and its variants
 - ✓ Mean exponential growth of nearby orbits (MEGNO)
 - ✓ Relative Lyapunov Indicator (RLI)
 - ✓ Dynamical Spectra

Lyapunov Exponents

Roughly speaking, the Lyapunov exponents of a given orbit characterize the **mean exponential rate of divergence** of trajectories surrounding it.

Consider an orbit in an M-dimensional phase space with **initial condition $x(0)$** and an **initial deviation vector from it $v(0)$** . Then the mean exponential rate of divergence is:

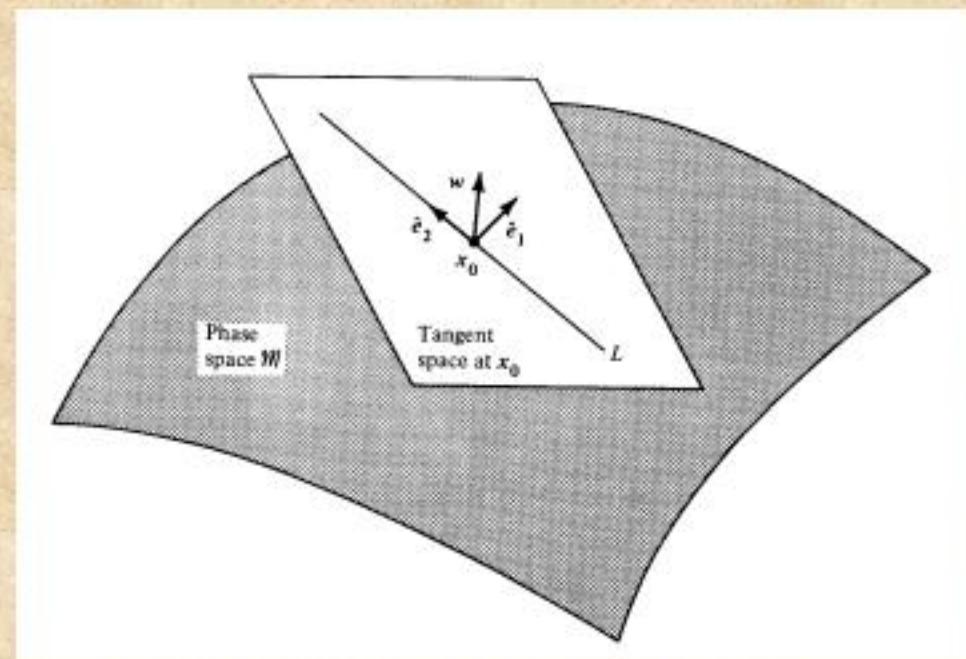
$$\sigma(x(0), v(0)) = \lim_{t \rightarrow \infty} \frac{1}{t} \ln \frac{\|v(t)\|}{\|v(0)\|}$$

Lyapunov Exponents

There exists an **M-dimensional basis $\{\hat{e}_i\}$** of v such that for any v , σ takes one of the M (possibly nondistinct) values

$$\sigma_i(x(0)) = \sigma(x(0), \hat{e}_i)$$

which are the **Lyapunov exponents.**



Benettin & Galgani, 1979, in Laval and Gressillon (eds.), op cit, 93

In autonomous Hamiltonian systems the M exponents are ordered in pairs of opposite sign numbers and two of them are 0.

Computation of the Maximal Lyapunov Exponent

Due to the exponential growth of $v(t)$ (and of $d(t) = ||v(t)||$) we **renormalize $v(t)$** from time to time.

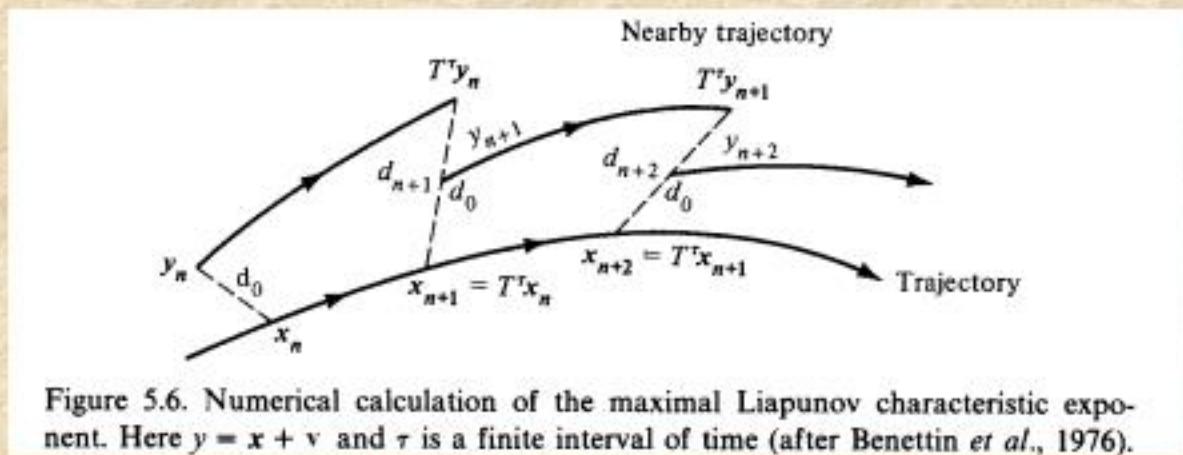


Figure 5.6. Numerical calculation of the maximal Liapunov characteristic exponent. Here $y = x + v$ and τ is a finite interval of time (after Benettin *et al.*, 1976).

Then the Maximal Lyapunov exponent is computed as

$$\sigma_1 = \lim_{n \rightarrow \infty} \frac{1}{n\tau} \sum_{i=1}^n \ln d_i$$

Maximal Lyapunov Exponent

$\sigma_1 = 0 \rightarrow$ Regular motion
 $\sigma_1 \neq 0 \rightarrow$ Chaotic motion

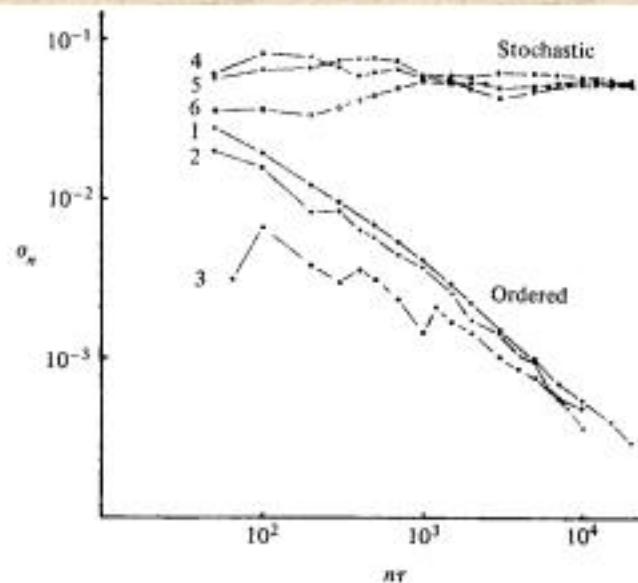


Figure 5.7. Behavior of σ_{\max} at the intermediate energy $E = 0.125$ for initial points taken in the ordered (curves 1–3) or stochastic (curves 4–6) regions (after Benettin et al., 1976).

Benettin et al. (1980) proposed an algorithm for the computation of all Lyapunov exponents.

Over the years several methods for the computation of few Lyapunov exponents have been developed (e.g. Greene & Kim 1987, Bridges & Reich 2001).

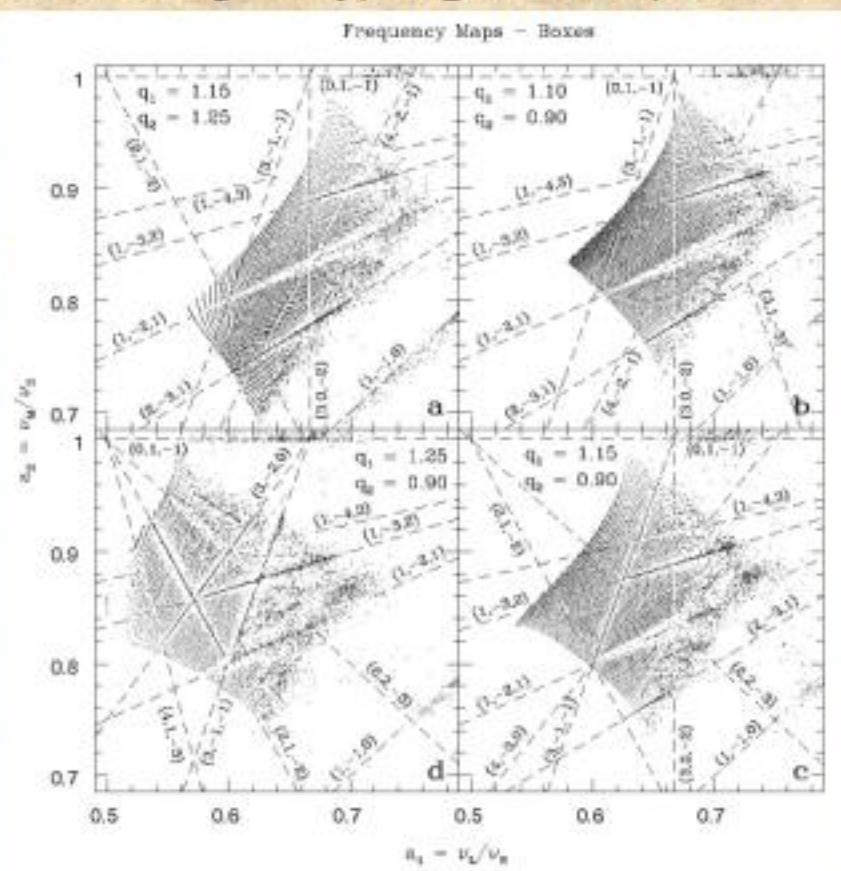
Frequency Analysis

We compute the **fundamental frequencies of an orbit** using the NAFF (Numerical Analysis of the Fundamental Frequency) algorithm (Laskar 1988).

Frequency Maps

Regular motion : The computed frequencies do not vary in time

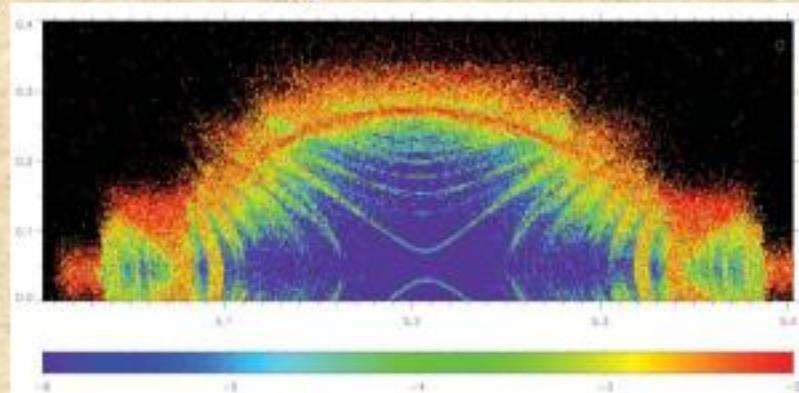
Chaotic motion : The computed frequencies vary in time



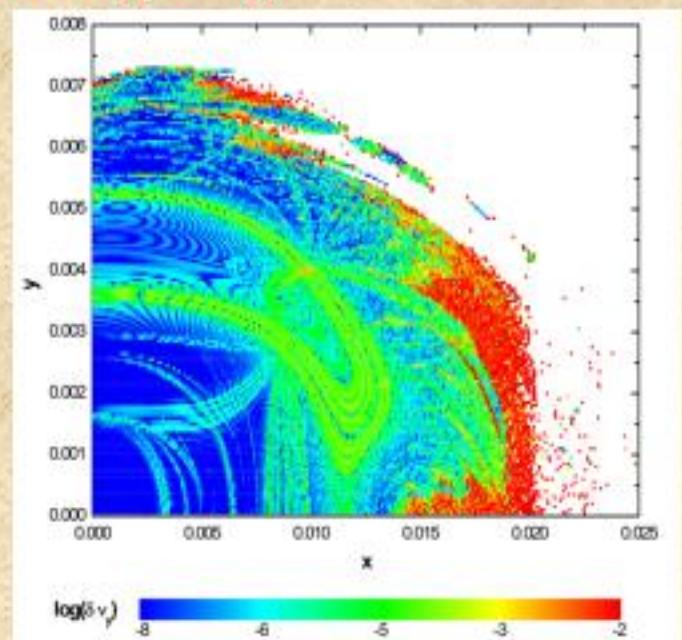
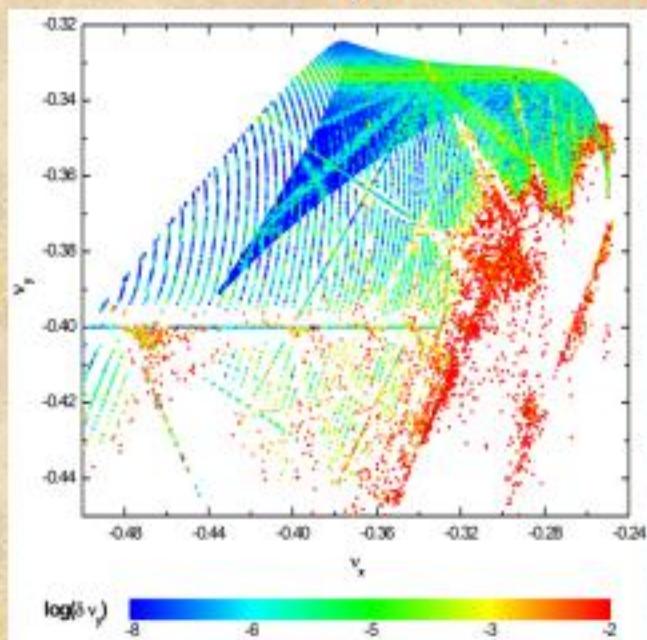
Papaphilippou Y and Laskar J 1998 Astron. Astrophys. 329 451

Frequency Analysis

Stability of Trojan asteroids, (a, e) diagram (Robutel & Gabern 2006)



Dynamics of ESRF storage ring



The Smaller ALignment Index (SALI) method

Definition of Smaller Alignment Index (SALI)

Consider the **n**-dimensional phase space of a conservative dynamical system (**symplectic map or Hamiltonian flow**).

An orbit in that space with initial condition :

$$P(0) = (x_1(0), x_2(0), \dots, x_n(0))$$

and a **deviation vector**

$$v(0) = (dx_1(0), dx_2(0), \dots, dx_n(0))$$

The evolution in time (in maps the time is discrete and is equal to the number **N** of the iterations) of a deviation vector is defined by:

- the **variational equations** (for Hamiltonian flows) and
- the **equations of the tangent map** (for mappings)

Definition of SALI

We follow the evolution in time of two different initial deviation vectors ($v_1(0)$, $v_2(0)$), and define SALI (Skokos, 2001, J. Phys. A) as:

$$\text{SALI}(t) = \min \left\{ \|\hat{v}_1(t) + \hat{v}_2(t)\|, \|\hat{v}_1(t) - \hat{v}_2(t)\| \right\}$$

where

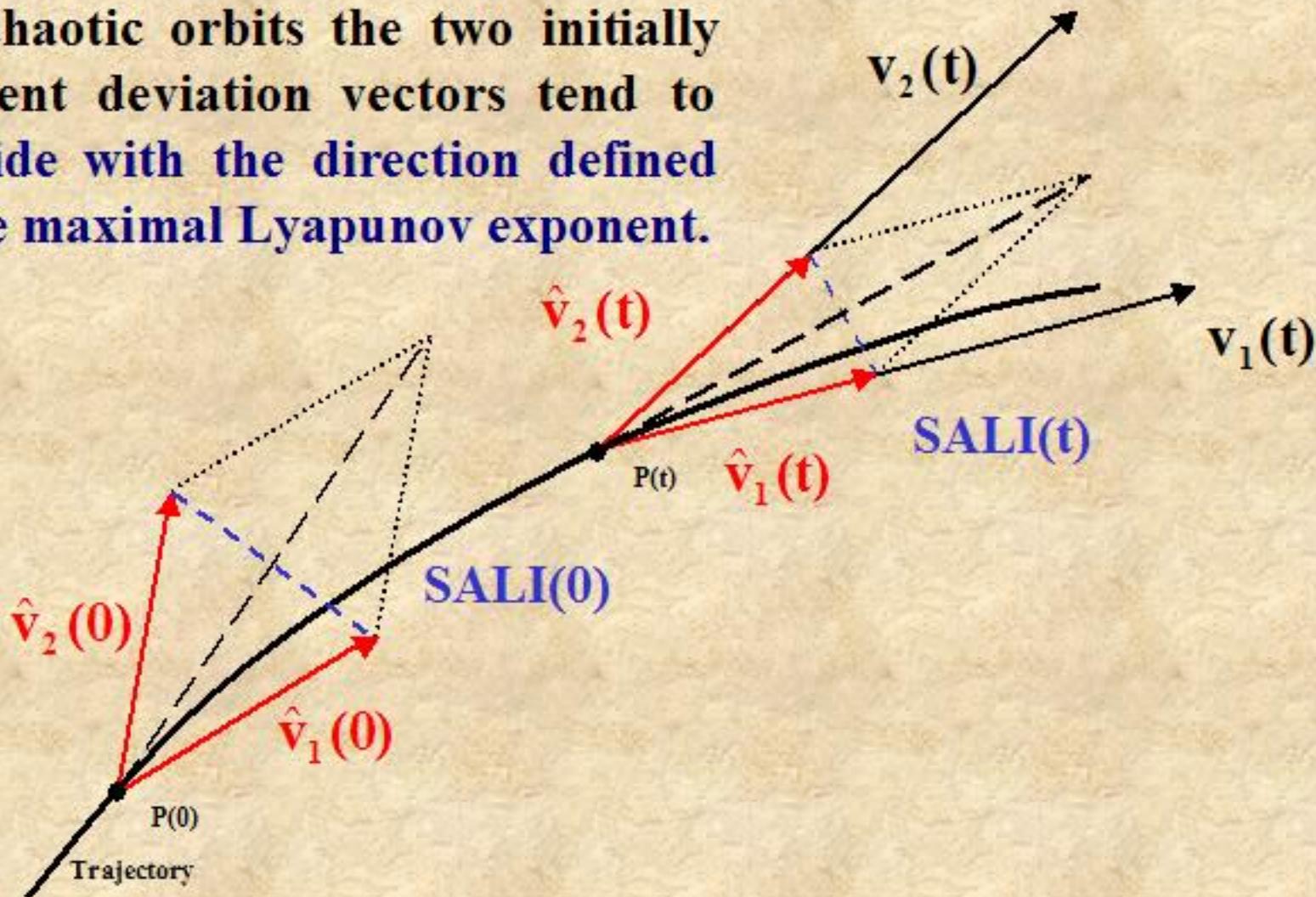
$$\hat{v}_1(t) = \frac{v_1(t)}{\|v_1(t)\|}$$

When the two vectors become **collinear**

$$\text{SALI}(t) \rightarrow 0$$

Behavior of SALI for chaotic motion

For chaotic orbits the two initially different deviation vectors tend to coincide with the direction defined by the maximal Lyapunov exponent.



Behavior of SALI for chaotic motion

The evolution of a deviation vector can be approximated by:

$$\mathbf{v}_1(t) = \sum_{i=1}^n c_i^{(1)} e^{\sigma_i t} \hat{\mathbf{u}}_i \approx c_1^{(1)} e^{\sigma_1 t} \hat{\mathbf{u}}_1 + c_2^{(1)} e^{\sigma_2 t} \hat{\mathbf{u}}_2$$

where $\sigma_1 > \sigma_2 \geq \dots \geq \sigma_n$ are the Lyapunov exponents.

In this approximation, we derive a leading order estimate of the ratio

$$\frac{\mathbf{v}_1(t)}{\|\mathbf{v}_1(t)\|} \approx \frac{c_1^{(1)} e^{\sigma_1 t} \hat{\mathbf{u}}_1 + c_2^{(1)} e^{\sigma_2 t} \hat{\mathbf{u}}_2}{|c_1^{(1)}| e^{\sigma_1 t}} = \pm \hat{\mathbf{u}}_1 + \frac{c_2^{(1)}}{|c_1^{(1)}|} e^{-(\sigma_1 - \sigma_2)t} \hat{\mathbf{u}}_2$$

and an analogous expression for \mathbf{v}_2

$$\frac{\mathbf{v}_2(t)}{\|\mathbf{v}_2(t)\|} \approx \frac{c_1^{(2)} e^{\sigma_1 t} \hat{\mathbf{u}}_1 + c_2^{(2)} e^{\sigma_2 t} \hat{\mathbf{u}}_2}{|c_1^{(2)}| e^{\sigma_1 t}} = \pm \hat{\mathbf{u}}_1 + \frac{c_2^{(2)}}{|c_1^{(2)}|} e^{-(\sigma_1 - \sigma_2)t} \hat{\mathbf{u}}_2$$

So we get:

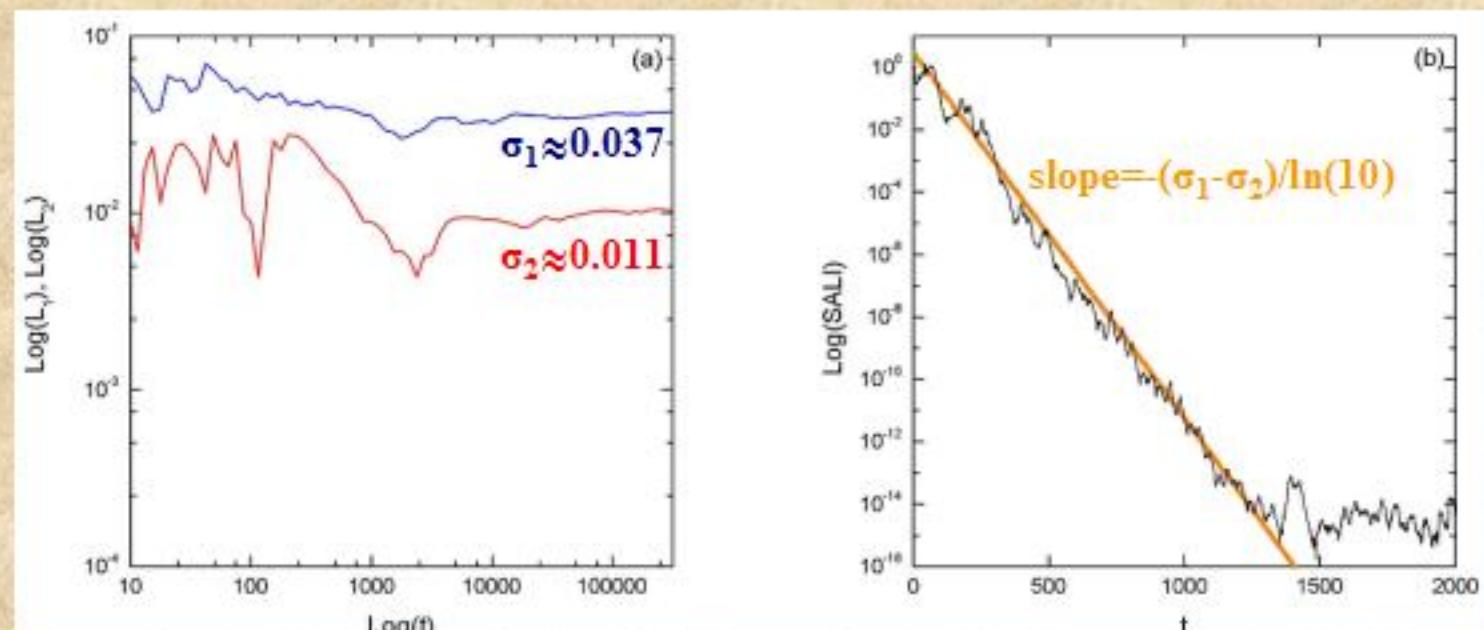
$$\text{SALI}(t) = \min \left\{ \left\| \frac{\mathbf{v}_1(t)}{\|\mathbf{v}_1(t)\|} + \frac{\mathbf{v}_2(t)}{\|\mathbf{v}_2(t)\|} \right\|, \left\| \frac{\mathbf{v}_1(t)}{\|\mathbf{v}_1(t)\|} - \frac{\mathbf{v}_2(t)}{\|\mathbf{v}_2(t)\|} \right\| \right\} \approx \left| \frac{c_2^{(1)}}{|c_1^{(1)}|} \pm \frac{c_2^{(2)}}{|c_1^{(2)}|} \right| e^{-(\sigma_1 - \sigma_2)t}$$

Behavior of SALI for chaotic motion

We test the validity of the approximation $\text{SALI} \propto e^{-(\sigma_1 - \sigma_2)t}$ (Skokos et al., 2004, J. Phys. A) for a chaotic orbit of the 3D Hamiltonian

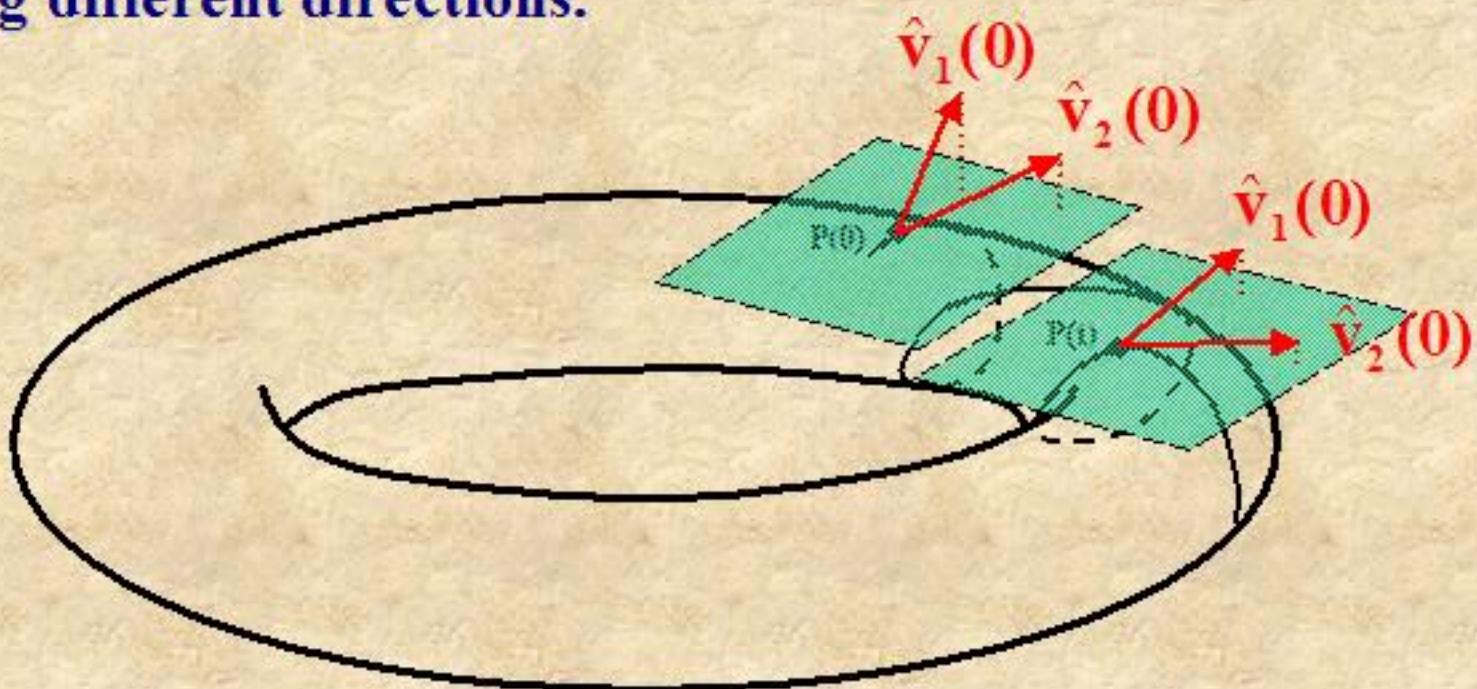
$$H = \sum_{i=1}^3 \frac{\omega_i}{2} (q_i^2 + p_i^2) + q_1^2 q_2 + q_1^2 q_3$$

with $\omega_1=1$, $\omega_2=1.4142$, $\omega_3=1.7321$, $H=0.09$



Behavior of SALI for regular motion

Regular motion occurs on a torus and two different initial deviation vectors become tangent to the torus, generally having different directions.



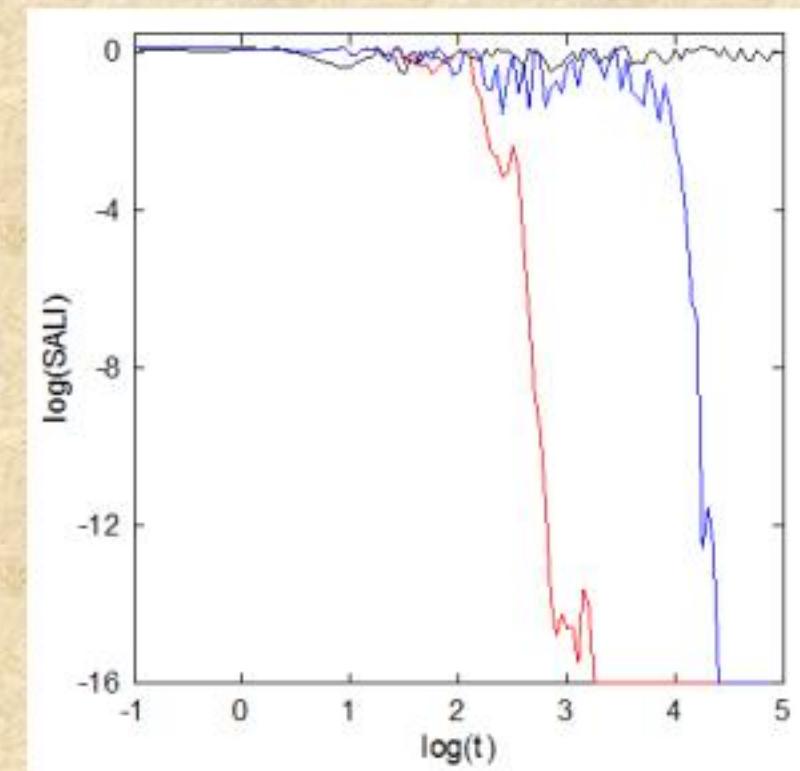
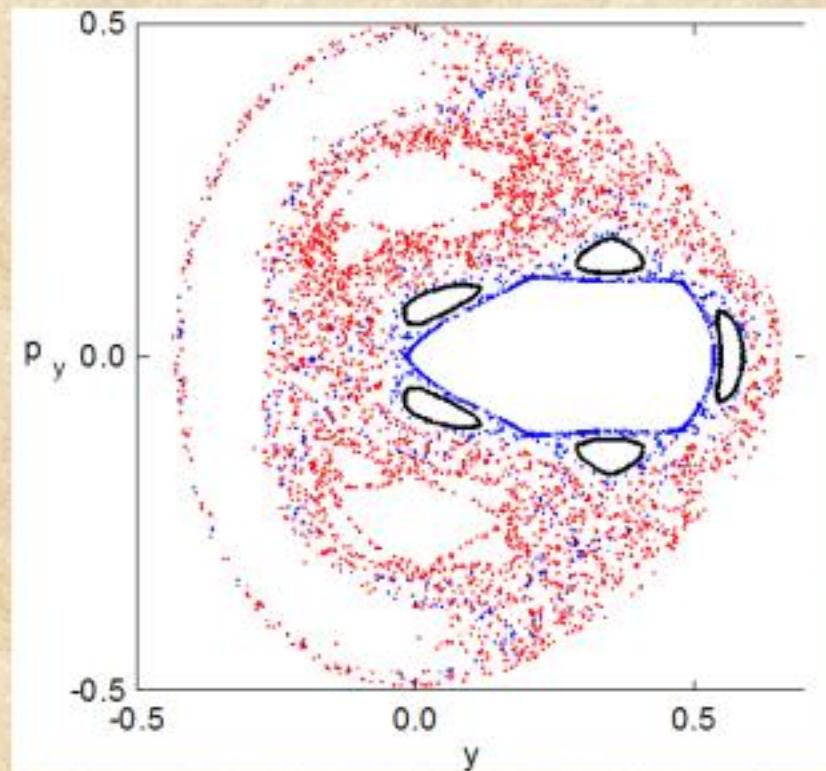
Applications – Hénon-Heiles system

For $E=1/8$ we consider the orbits with initial conditions:

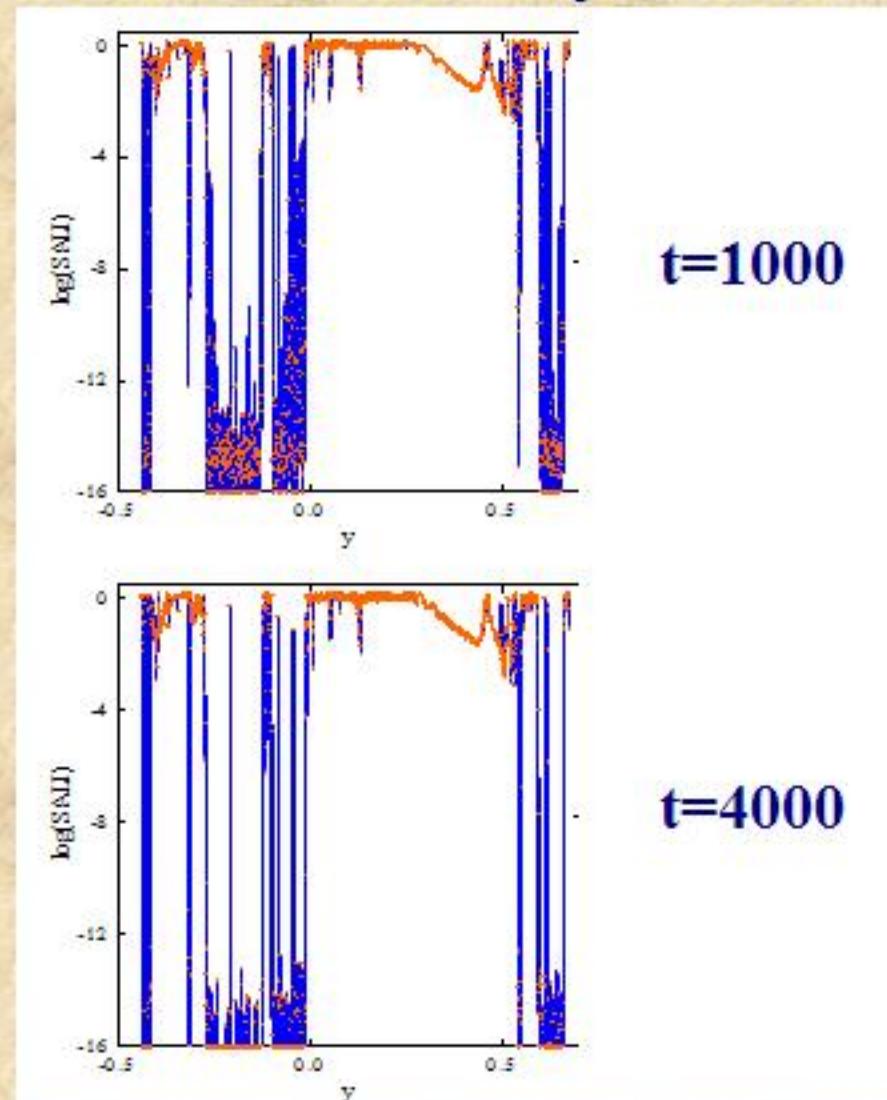
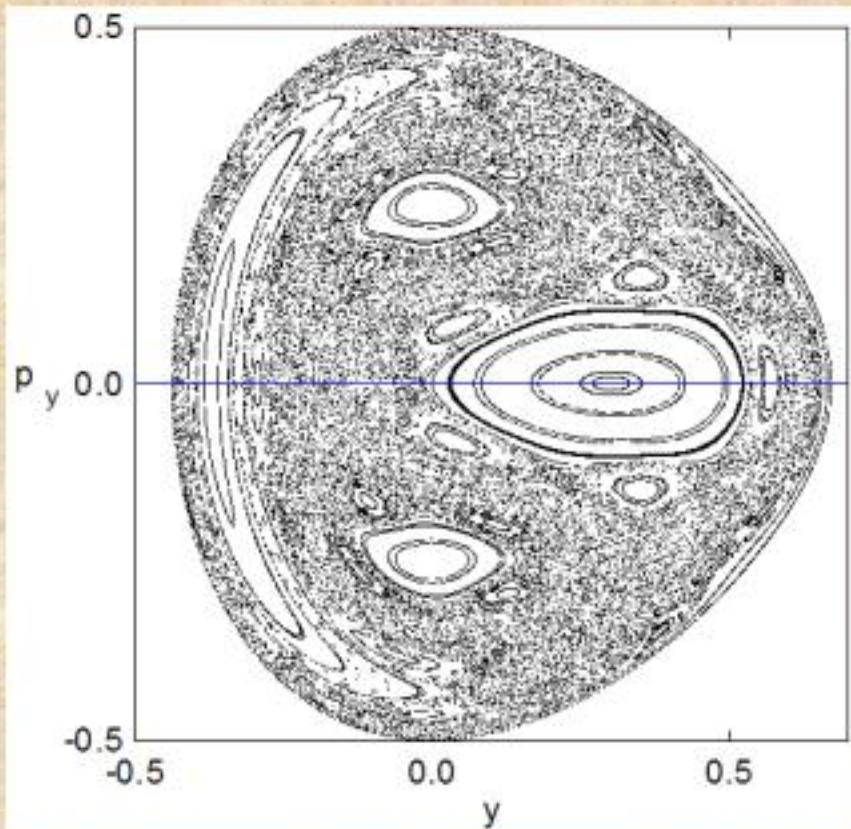
Ordered orbit, $x=0$, $y=0.55$, $p_x=0.2417$, $p_y=0$

Chaotic orbit, $x=0$, $y=-0.016$, $p_x=0.49974$, $p_y=0$

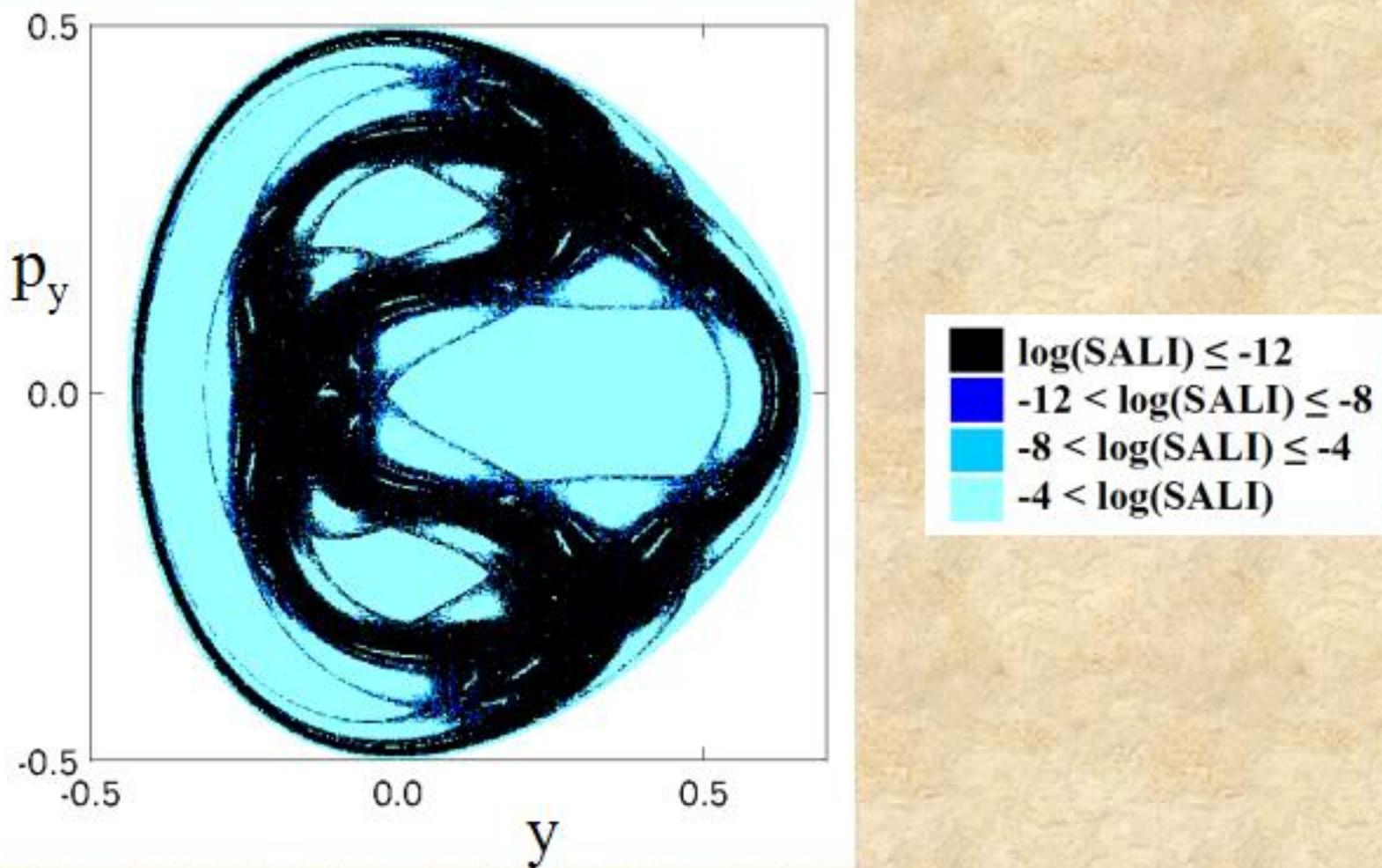
Chaotic orbit, $x=0$, $y=-0.01344$, $p_x=0.49982$, $p_y=0$



Applications – Hénon-Heiles system

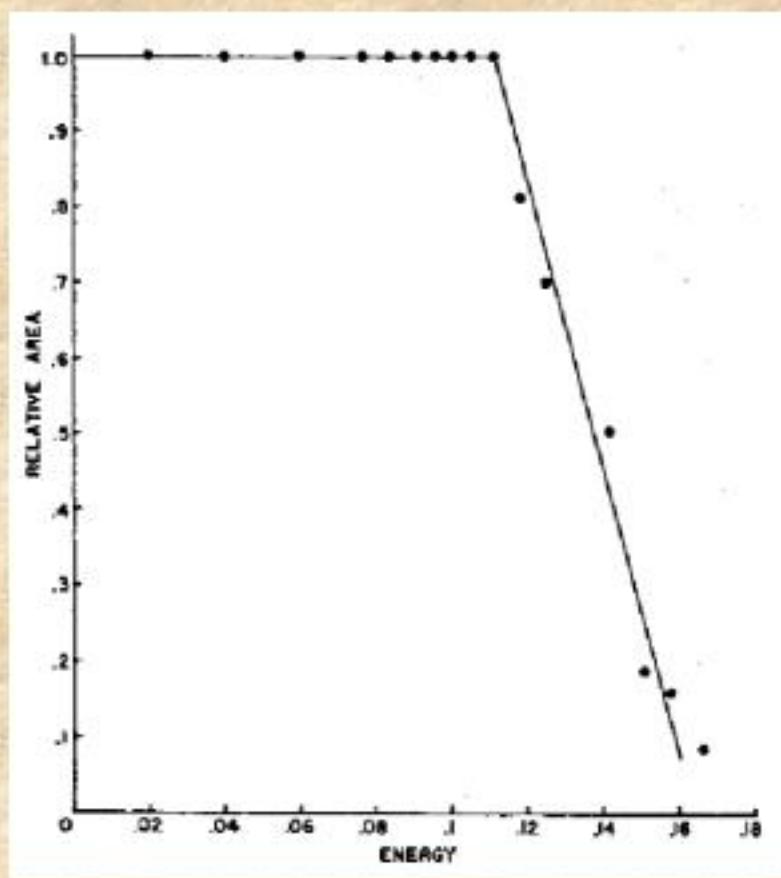


Applications – Hénon-Heiles system

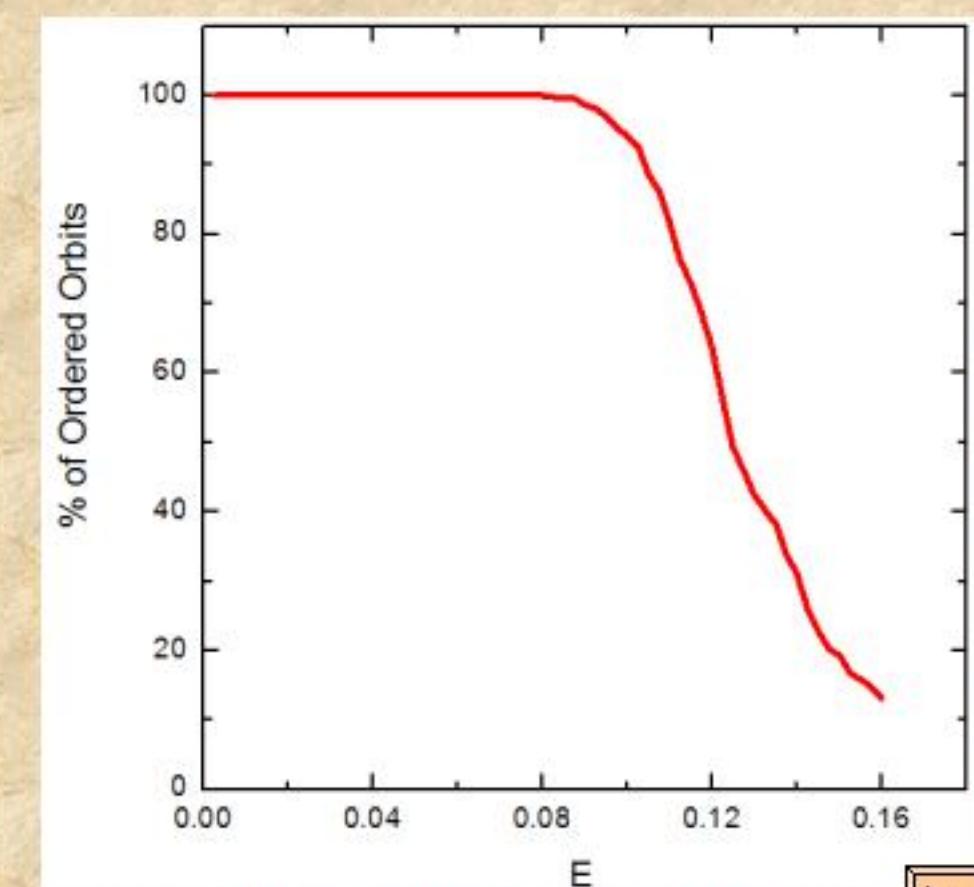


Applications – Hénon-Heiles system

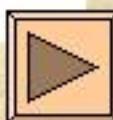
The percentage of non chaotic orbits ($\text{SALI} > 10^{-8}$ for $t=1000$)



Hénon-Heiles (1964) Astron. J. 69, 73.



A. Manos (2004) Master Thesis, Univ. of Patras



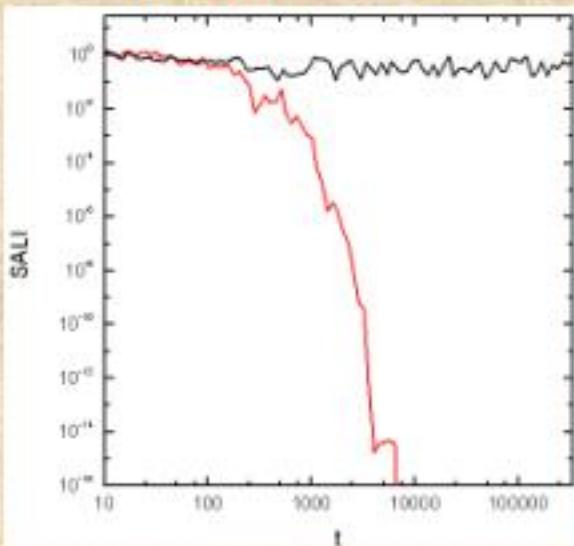
Applications – 3D Hamiltonian

We consider the 3D Hamiltonian

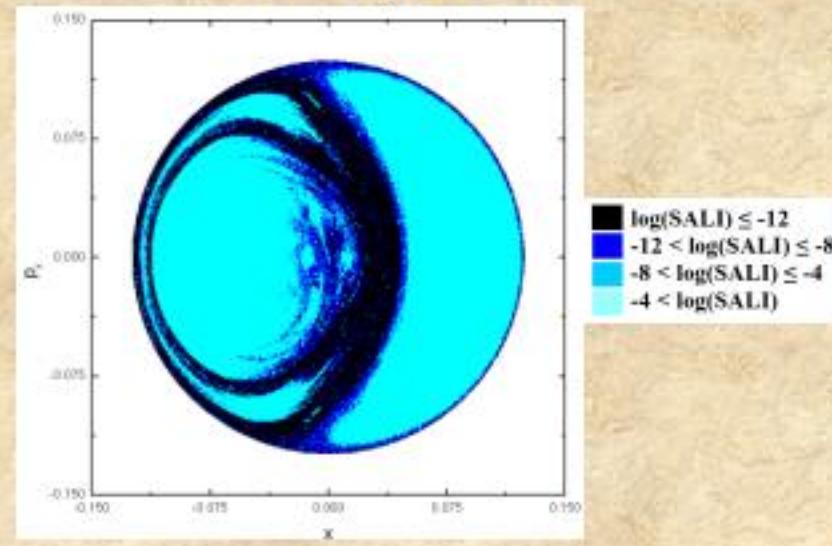
$$H = \frac{1}{2}(p_x^2 + p_y^2 + p_z^2) + \frac{1}{2}(Ax^2 + By^2 + Cz^2) - \varepsilon xz^2 - \eta yz^2$$

with $A=0.9$, $B=0.4$, $C=0.225$, $\varepsilon=0.56$, $\eta=0.2$, $H=0.00765$.

Behavior of the SALI for ordered
and **chaotic** orbits



Color plot of the subspace
 $y=z=p_y=0$, $p_z > 0$



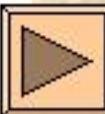
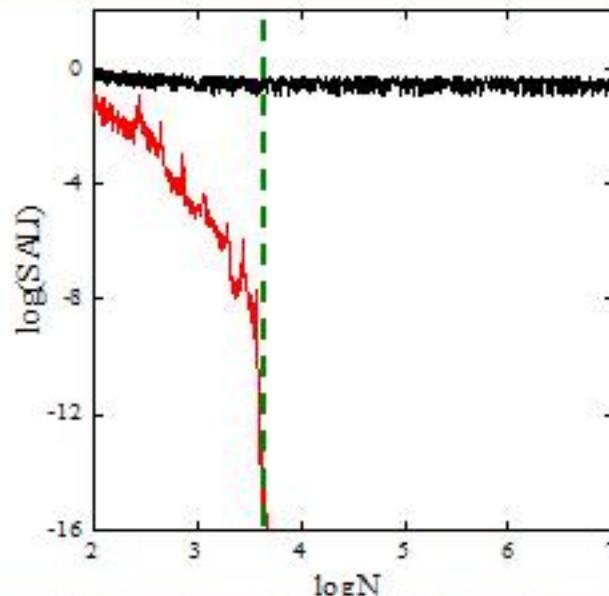
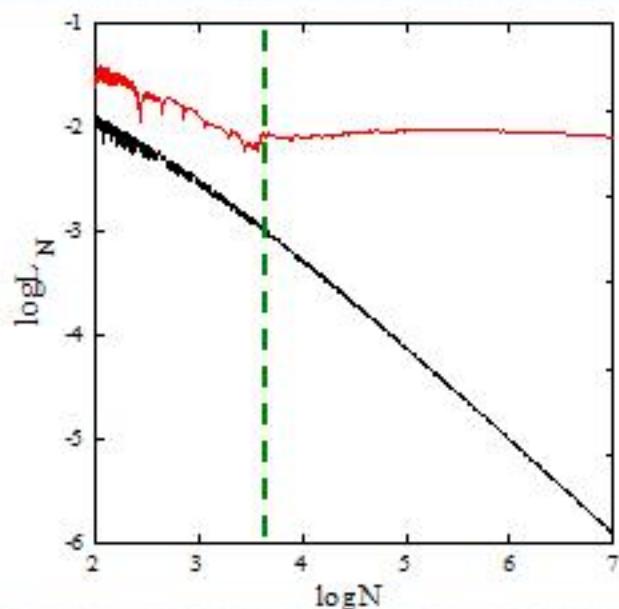
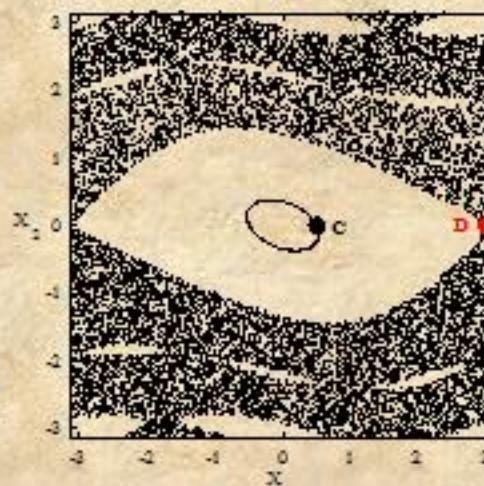
Applications – 4D map

$$\begin{aligned}x'_1 &= x_1 + x_2 \\x'_2 &= x_2 - \nu \sin(x_1 + x_2) - \mu [1 - \cos(x_1 + x_2 + x_3 + x_4)] \quad (\text{mod } 2\pi) \\x'_3 &= x_3 + x_4 \\x'_4 &= x_4 - \kappa \sin(x_3 + x_4) - \mu [1 - \cos(x_1 + x_2 + x_3 + x_4)]\end{aligned}$$

For $\nu=0.5$, $\kappa=0.1$, $\mu=0.1$ we consider the orbits:

ordered orbit C with initial conditions $x_1=0.5$, $x_2=0$, $x_3=0.5$, $x_4=0$.

chaotic orbit D with initial conditions $x_1=3$, $x_2=0$, $x_3=0.5$, $x_4=0$.



Applications – 4D Accelerator map

We consider the 4D symplectic map

$$\begin{pmatrix} x'_1 \\ x'_2 \\ x'_3 \\ x'_4 \end{pmatrix} = \begin{pmatrix} \cos\omega_1 & -\sin\omega_1 & 0 & 0 \\ \sin\omega_1 & \cos\omega_1 & 0 & 0 \\ 0 & 0 & \cos\omega_2 & -\sin\omega_2 \\ 0 & 0 & \sin\omega_2 & \cos\omega_2 \end{pmatrix} \times \begin{pmatrix} x_1 \\ x_2 + x_1^2 - x_3^2 \\ x_3 \\ x_4 - 2x_1x_3 \end{pmatrix}$$

describing the **instantaneous sextupole ‘kicks’** experienced by a particle as it passes through an accelerator (Turchetti & Scandale 1991, Bountis & Tompaidis 1991, Vrahatis et al. 1996, 1997).

x_1 and x_3 are the particle’s deflections from the ideal circular orbit, in the horizontal and vertical directions respectively.

x_2 and x_4 are the associated momenta

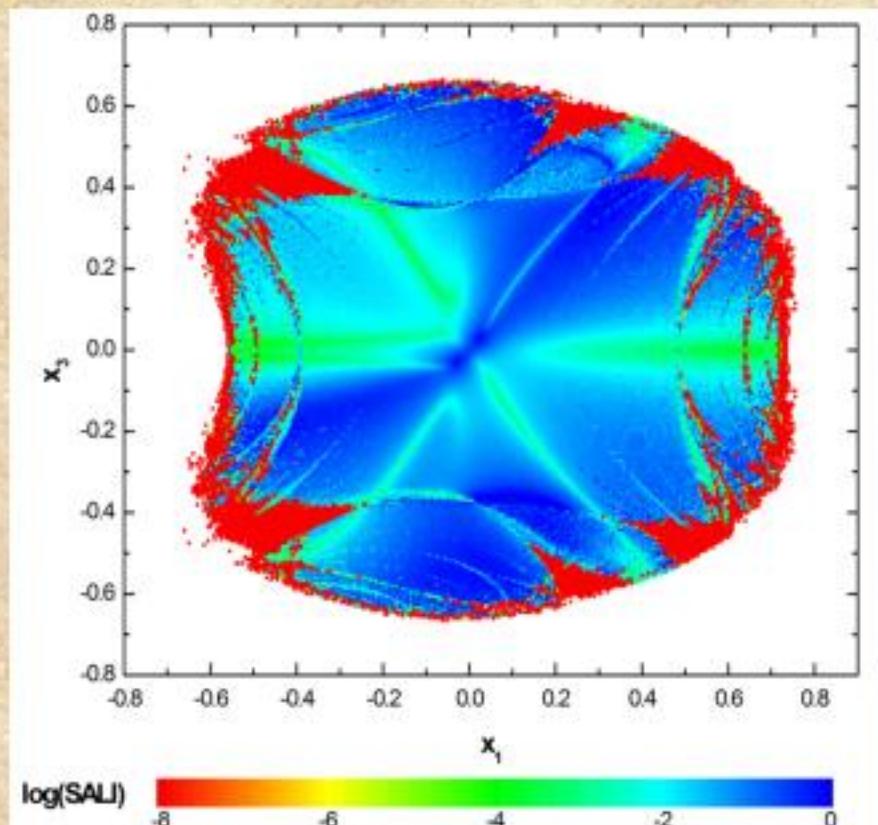
ω_1, ω_2 are related to the accelerator’s tunes q_x, q_y by

$$\omega_1 = 2\pi q_x, \quad \omega_2 = 2\pi q_y$$

Our problem is to estimate the **region of stability** of the particle’s motion, the so-called **dynamic aperture** of the beam (Bountis & Skokos, 2006, Nucl. Inst Meth. Phys Res. A).

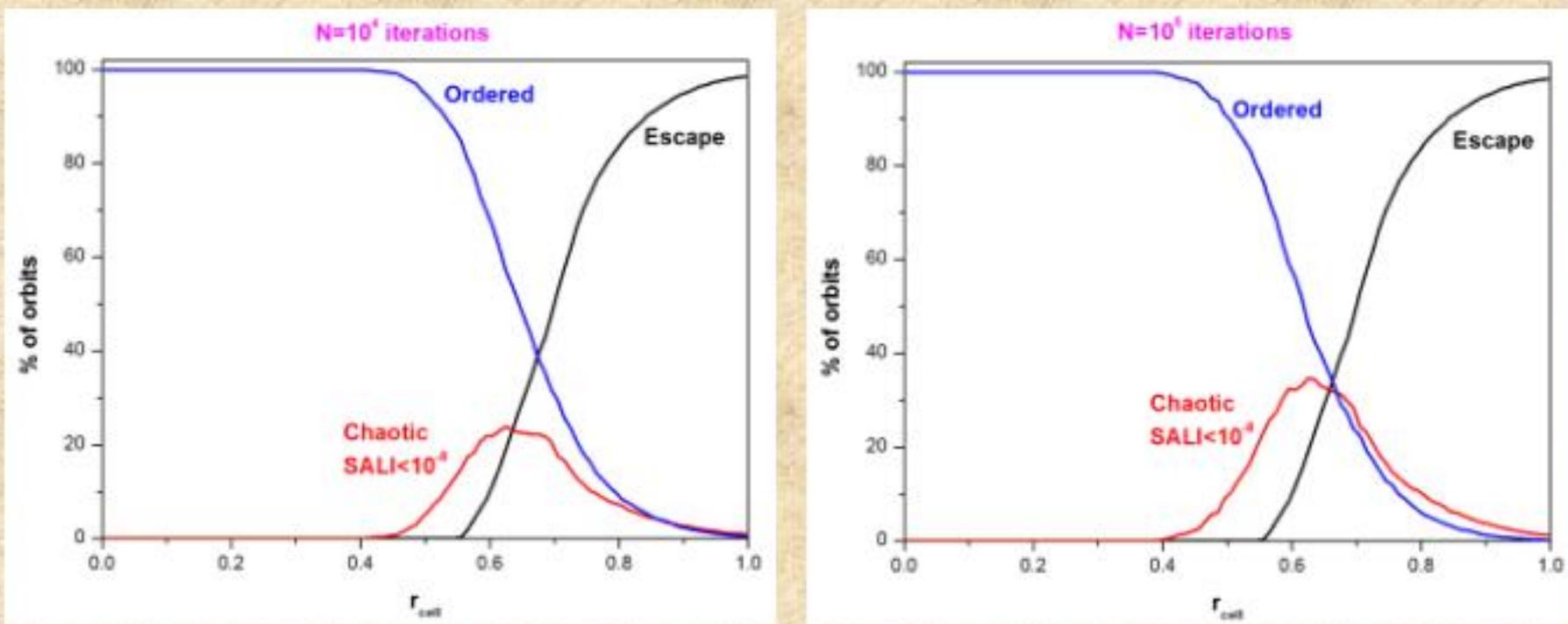
4D Accelerator map – "Global" study

Regions of different values of the SALI on the subspace
 $x_2(0)=x_4(0)=0$, after 10^4 iterations ($q_x=0.61803$ $q_y=0.4152$)



4D Accelerator map – "Global" study

We consider 1,922,833 orbits by varying all x_1, x_2, x_3, x_4 within spherical shells of width 0.01 in a hypersphere of radius 1. ($q_x = 0.61803$ $q_y = 0.4152$)



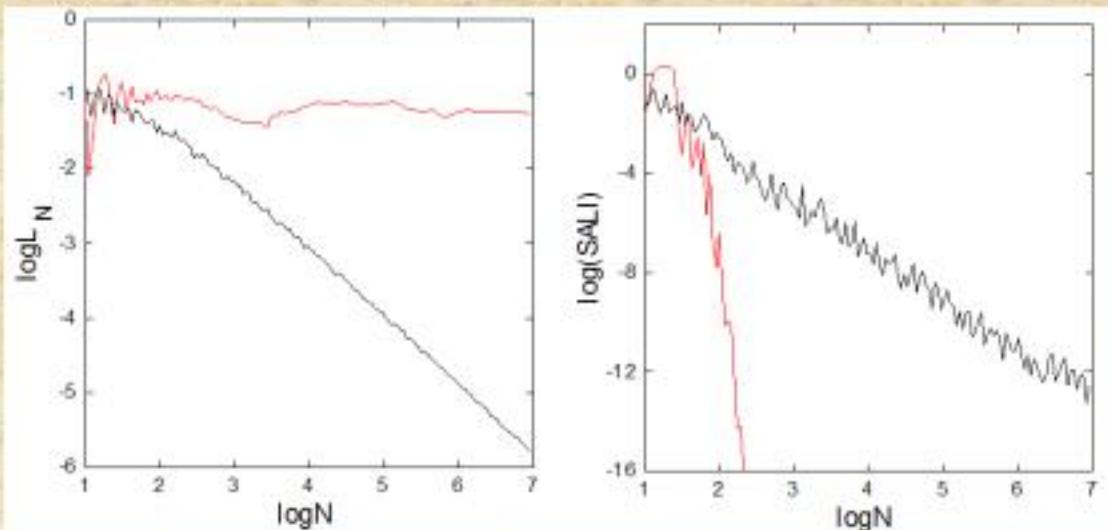
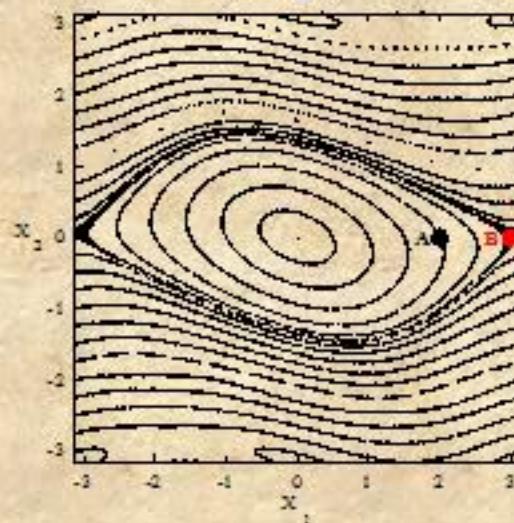
Applications – 2D map

$$\begin{aligned}x'_1 &= x_1 + x_2 \\x'_2 &= x_2 - v \sin(x_1 + x_2)\end{aligned}\quad (\text{mod } 2\pi)$$

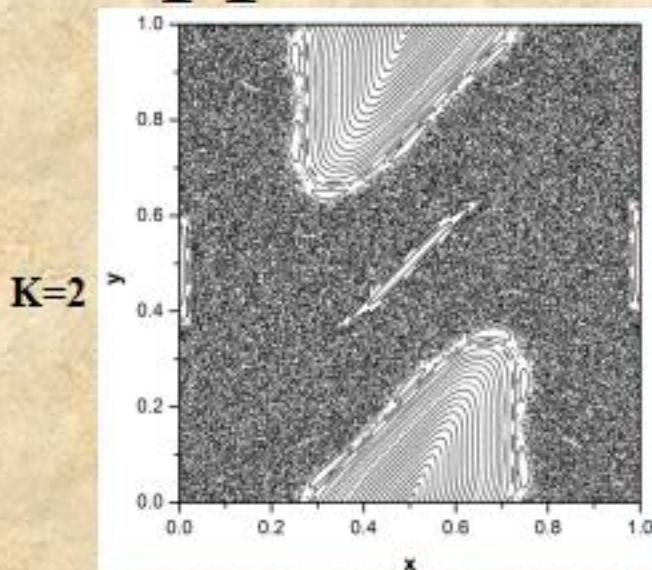
For $v=0.5$ we consider the orbits:

ordered orbit A with initial conditions $x_1=2, x_2=0$.

chaotic orbit B with initial conditions $x_1=3, x_2=0$.



Applications – Standard map

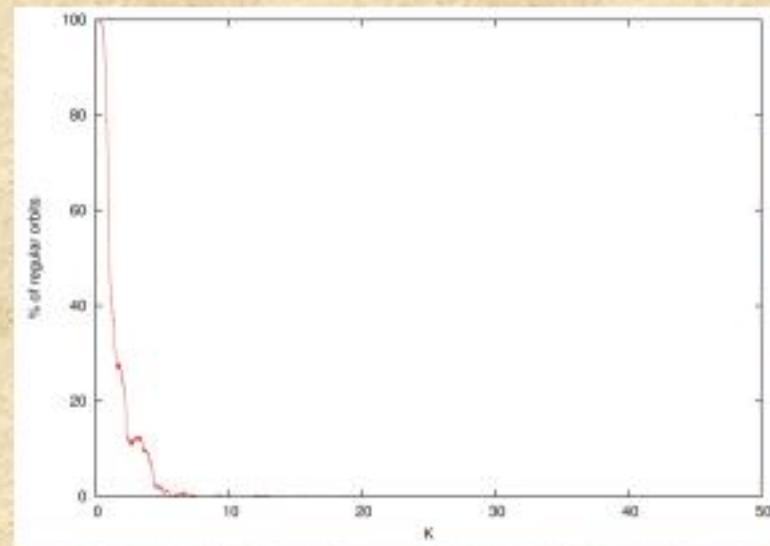


$$\begin{aligned}x' &= x + y' \\y' &= y + \frac{K}{2\pi} \sin(2\pi x) \quad (\text{mod } 1)\end{aligned}$$

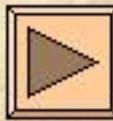
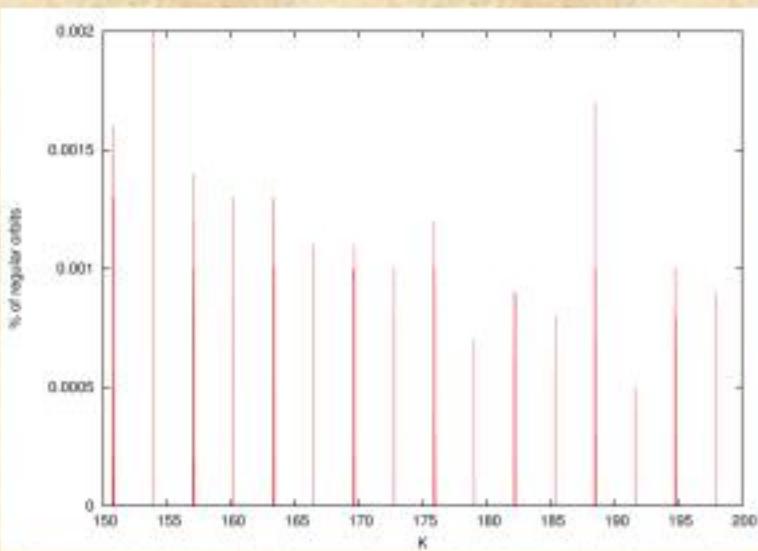
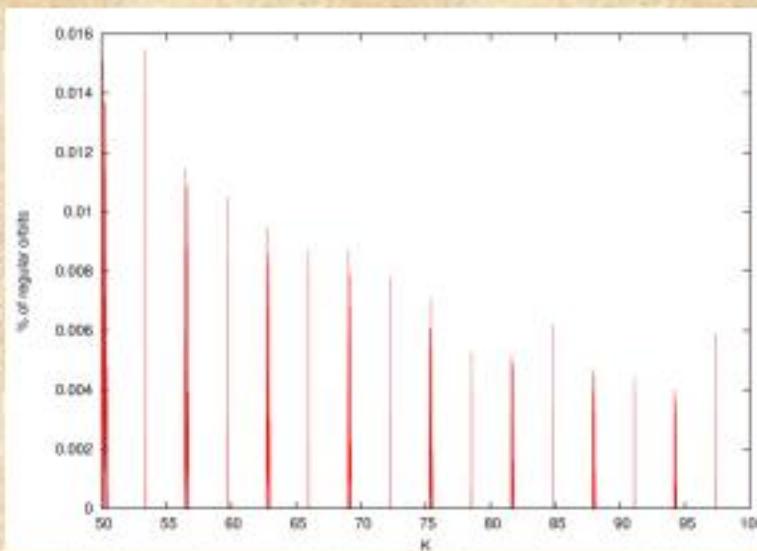
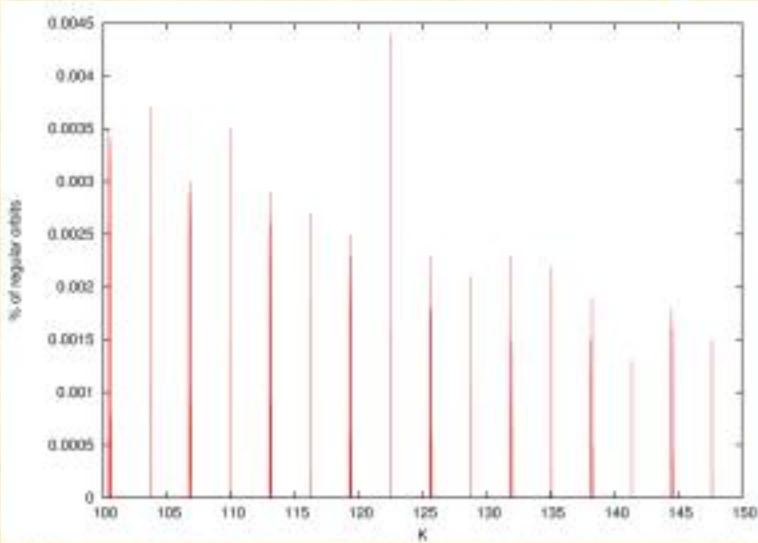
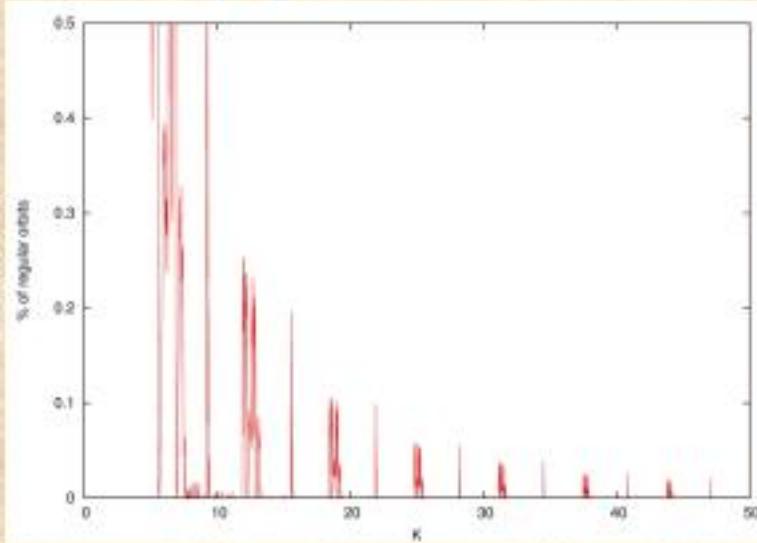
Dependence of the percentage of regular motion on K
(Dvorak et al. 2003).

For every K ($dK=0.001$) we use a **1000 X 1000 grid** of initial conditions. $N=500$.

Chaotic orbits: $\text{SALI} < 10^{-10}$

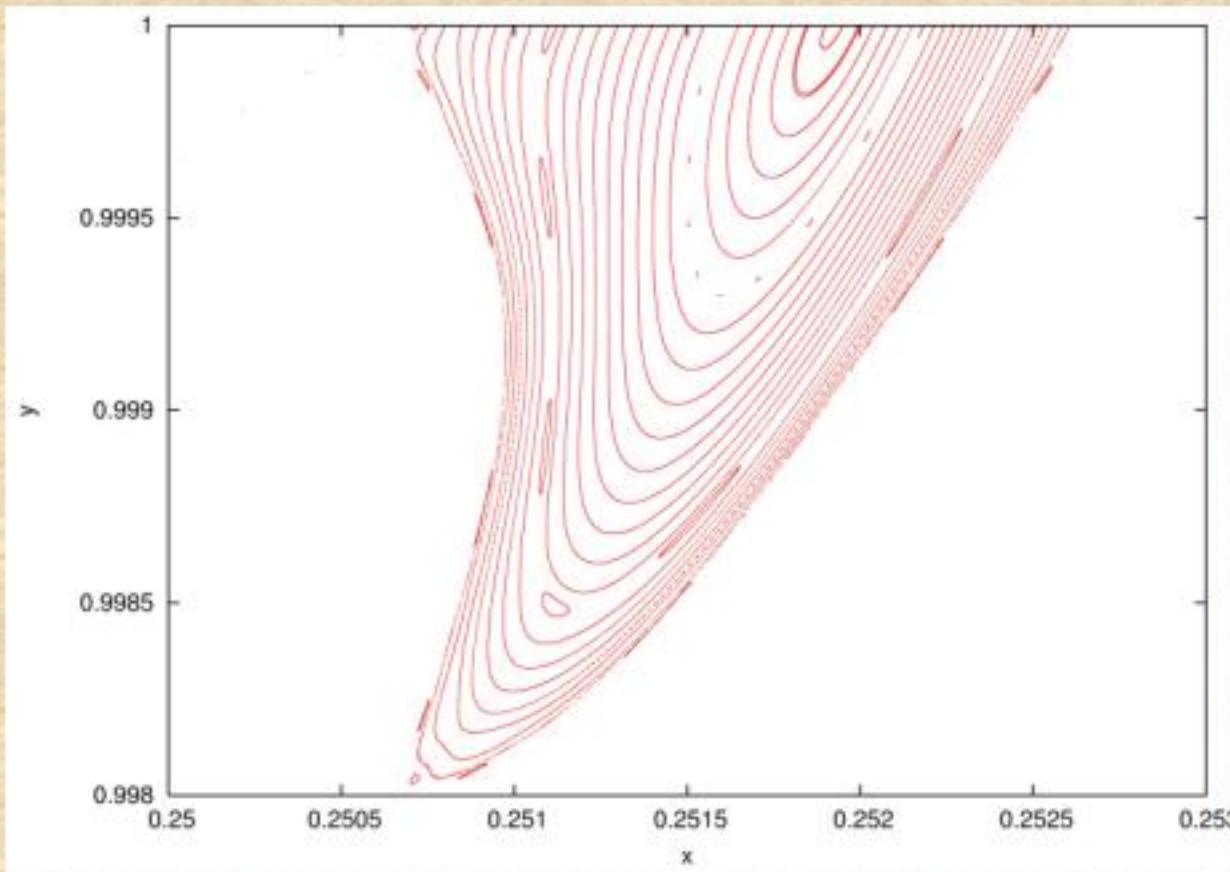


Applications – Standard map



Applications – Standard map

For $K=194.793$ we have only 0.0004% regular orbits.



Behavior of SALI

2D maps

SALI $\rightarrow 0$ both for regular and chaotic orbits

following, however, completely different time rates which allows us to distinguish between the two cases.

Hamiltonian flows and multidimensional maps

SALI $\rightarrow 0$ for chaotic orbits

SALI \rightarrow constant $\neq 0$ for regular orbits

Questions

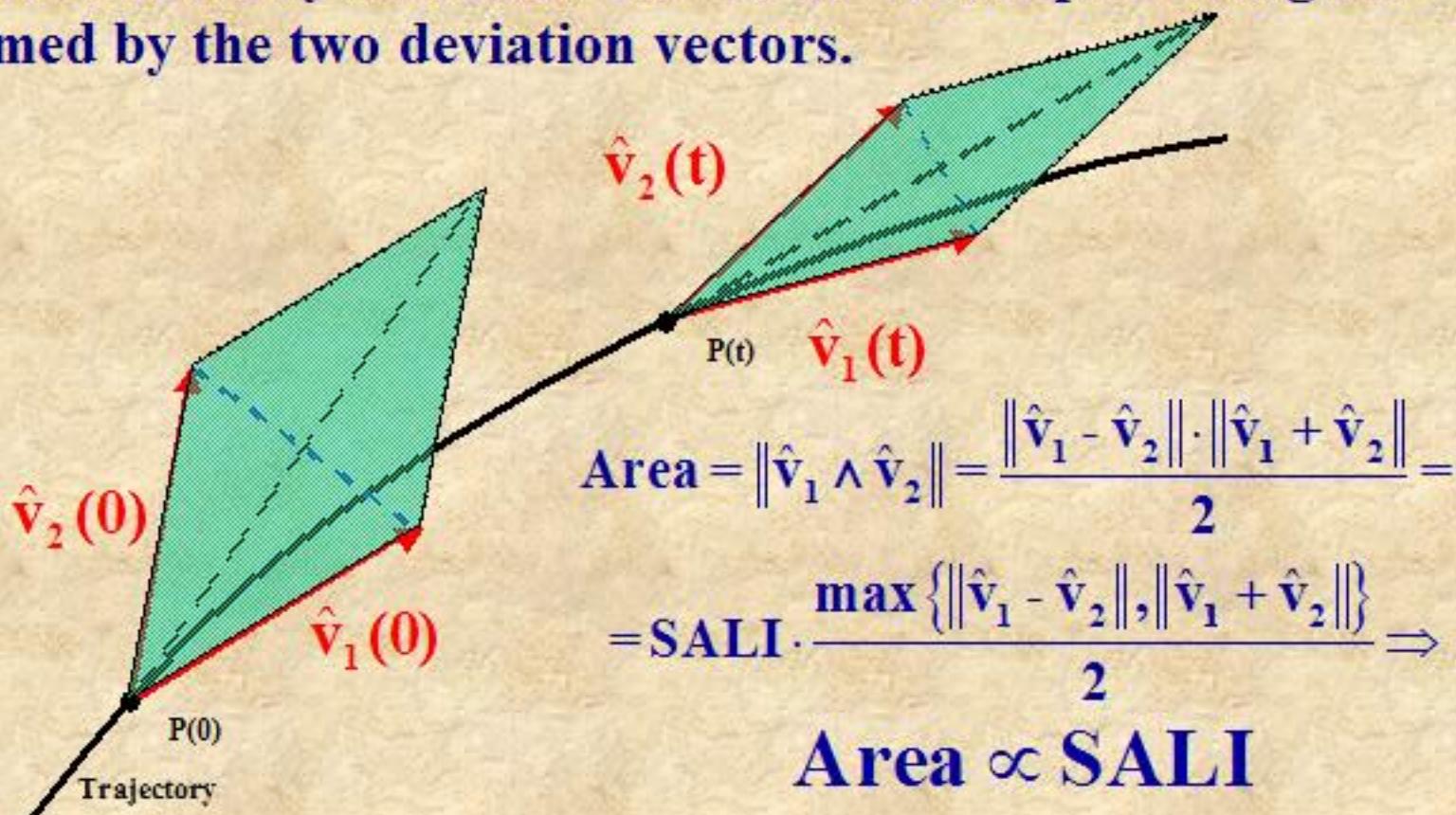
Can we generalize SALI so that the new index:

- Can rapidly reveal the nature of chaotic orbits with $\sigma_1 \approx \sigma_2$ ($SALI \propto e^{-(\sigma_1 - \sigma_2)t}$)?
- Depends on several Lyapunov exponents for chaotic orbits?
- Exhibits power-law decay for regular orbits depending on the dimensionality of the tangent space of the reference orbit as for 2D maps?

The Generalized ALignment Indices (GALIs) method

Definition of Generalized Alignment Index (GALI)

SALI effectively measures the ‘area’ of the parallelogram formed by the two deviation vectors.



Definition of GALI

Generalizing this idea to an N degree of freedom Hamiltonian system (Skokos et al., 2006, submitted), we follow the evolution of

k deviation vectors with $2 \leq k \leq 2N$,

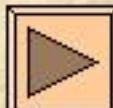
and

define the Generalized Alignment Index (GALI) of order k :

$$\text{GALI}_k(t) = \|\hat{v}_1(t) \wedge \hat{v}_2(t) \wedge \dots \wedge \hat{v}_k(t)\|$$

Clearly:

$$\text{GALI}_2(t) \propto \text{SALI}(t)$$



Computation of wedge product

We consider as a basis of the $2N$ -dimensional tangent space of the Hamiltonian flow the usual set of orthonormal vectors:

$$\hat{\mathbf{e}}_1 = (1, 0, 0, \dots, 0), \hat{\mathbf{e}}_2 = (0, 1, 0, \dots, 0), \dots, \hat{\mathbf{e}}_{2N} = (0, 0, 0, \dots, 1)$$

Then for k deviation vectors we have:

$$\begin{bmatrix} \hat{\mathbf{v}}_1 \\ \hat{\mathbf{v}}_2 \\ \vdots \\ \hat{\mathbf{v}}_k \end{bmatrix} = \begin{bmatrix} \mathbf{v}_{11} & \mathbf{v}_{12} & \cdots & \mathbf{v}_{12N} \\ \mathbf{v}_{21} & \mathbf{v}_{22} & \cdots & \mathbf{v}_{22N} \\ \vdots & \vdots & \ddots & \vdots \\ \mathbf{v}_{k1} & \mathbf{v}_{k2} & \cdots & \mathbf{v}_{k2N} \end{bmatrix} \cdot \begin{bmatrix} \hat{\mathbf{e}}_1 \\ \hat{\mathbf{e}}_2 \\ \vdots \\ \hat{\mathbf{e}}_{2N} \end{bmatrix}$$

$$\hat{\mathbf{v}}_1 \wedge \hat{\mathbf{v}}_2 \wedge \cdots \wedge \hat{\mathbf{v}}_k = \sum_{1 \leq i_1 < i_2 < \cdots < i_k \leq 2N} \begin{bmatrix} \mathbf{v}_{1i_1} & \mathbf{v}_{1i_2} & \cdots & \mathbf{v}_{1i_k} \\ \mathbf{v}_{2i_1} & \mathbf{v}_{2i_2} & \cdots & \mathbf{v}_{2i_k} \\ \vdots & \vdots & \ddots & \vdots \\ \mathbf{v}_{ki_1} & \mathbf{v}_{ki_2} & \cdots & \mathbf{v}_{ki_k} \end{bmatrix} \hat{\mathbf{e}}_{i_1} \wedge \hat{\mathbf{e}}_{i_2} \wedge \cdots \wedge \hat{\mathbf{e}}_{i_k}$$

Computation of wedge product

We define as ‘norm’ of the wedge product the quantity :

$$\|\hat{\mathbf{v}}_1 \wedge \hat{\mathbf{v}}_2 \wedge \cdots \wedge \hat{\mathbf{v}}_k\| = \left\{ \sum_{1 \leq i_1 < i_2 < \cdots < i_k \leq 2N} \begin{vmatrix} \mathbf{v}_{1i_1} & \mathbf{v}_{1i_2} & \cdots & \mathbf{v}_{1i_k} \\ \mathbf{v}_{2i_1} & \mathbf{v}_{2i_2} & \cdots & \mathbf{v}_{2i_k} \\ \vdots & \vdots & \ddots & \vdots \\ \mathbf{v}_{ki_1} & \mathbf{v}_{ki_2} & \cdots & \mathbf{v}_{ki_k} \end{vmatrix}^2 \right\}^{1/2}$$

Computation of GALI - Example

Let us compute $GALI_3$ in the case of 2D Hamiltonian system (4-dimensional phase space).

$$\begin{bmatrix} \hat{\mathbf{v}}_1 \\ \hat{\mathbf{v}}_2 \\ \hat{\mathbf{v}}_3 \end{bmatrix} = \begin{bmatrix} \mathbf{v}_{11} & \mathbf{v}_{12} & \mathbf{v}_{13} & \mathbf{v}_{14} \\ \mathbf{v}_{21} & \mathbf{v}_{22} & \mathbf{v}_{23} & \mathbf{v}_{24} \\ \mathbf{v}_{31} & \mathbf{v}_{32} & \mathbf{v}_{33} & \mathbf{v}_{34} \end{bmatrix} \cdot \begin{bmatrix} \hat{\mathbf{e}}_1 \\ \hat{\mathbf{e}}_2 \\ \hat{\mathbf{e}}_3 \\ \hat{\mathbf{e}}_4 \end{bmatrix}$$

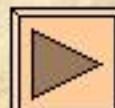
$$\text{GALI}_3 = \left\| \hat{\mathbf{v}}_1 \wedge \hat{\mathbf{v}}_2 \wedge \hat{\mathbf{v}}_3 \right\| = \sqrt{\left(v_{11}^2 + v_{12}^2 + v_{13}^2 \right)^2 + \left(v_{21}^2 + v_{22}^2 + v_{23}^2 \right)^2 + \left(v_{31}^2 + v_{32}^2 + v_{33}^2 \right)^2 + \left(v_{11}^2 + v_{12}^2 + v_{14}^2 \right)^2 + \left(v_{21}^2 + v_{22}^2 + v_{24}^2 \right)^2 + \left(v_{31}^2 + v_{32}^2 + v_{34}^2 \right)^2 + \left(v_{11}^2 + v_{13}^2 + v_{14}^2 \right)^2 + \left(v_{21}^2 + v_{23}^2 + v_{24}^2 \right)^2 + \left(v_{31}^2 + v_{33}^2 + v_{34}^2 \right)^2 + \left(v_{12}^2 + v_{13}^2 + v_{14}^2 \right)^2 + \left(v_{22}^2 + v_{23}^2 + v_{24}^2 \right)^2 + \left(v_{32}^2 + v_{33}^2 + v_{34}^2 \right)^2}^{1/2}$$

Behavior of GALI_k for chaotic motion

GALI_k ($2 \leq k \leq 2N$) tends exponentially to zero with exponents that involve the values of the first k largest Lyapunov exponents $\sigma_1, \sigma_2, \dots, \sigma_k$ (Skokos et al., 2006, submitted) :

$$\text{GALI}_k(t) \propto e^{-[(\sigma_1 - \sigma_2) + (\sigma_1 - \sigma_3) + \dots + (\sigma_1 - \sigma_k)]t}$$

The above relation is valid even if some Lyapunov exponents are equal, or very close to each other.



Behavior of GALI_k for chaotic motion

Using the approximation:

$$\mathbf{v}_i(t) = \sum_{j=1}^{2N} c_j^i e^{\sigma_j t} \hat{\mathbf{u}}_j = c_1^i e^{\sigma_1 t} \hat{\mathbf{u}}_1 + c_2^i e^{\sigma_2 t} \hat{\mathbf{u}}_2 + \dots + c_{2N}^i e^{\sigma_{2N} t} \hat{\mathbf{u}}_{2N}, \quad \|\mathbf{v}_i(t)\| \approx |c_1^i| e^{\sigma_1 t}$$

where $\sigma_1 \geq \sigma_2 \geq \dots \geq \sigma_n$ are the **Lyapunov exponents**, and $\hat{\mathbf{u}}_j$ $j=1, 2, \dots, 2N$ the corresponding eigendirections, we get

$$\begin{bmatrix} \hat{\mathbf{v}}_1 \\ \hat{\mathbf{v}}_2 \\ \vdots \\ \hat{\mathbf{v}}_k \end{bmatrix} = \begin{bmatrix} s_1 & \frac{c_2^1}{|c_1^1|} e^{-(\sigma_1 - \sigma_2)t} & \frac{c_3^1}{|c_1^1|} e^{-(\sigma_1 - \sigma_3)t} & \dots & \frac{c_{2N}^1}{|c_1^1|} e^{-(\sigma_1 - \sigma_{2N})t} \\ s_2 & \frac{c_2^2}{|c_1^2|} e^{-(\sigma_1 - \sigma_2)t} & \frac{c_3^2}{|c_1^2|} e^{-(\sigma_1 - \sigma_3)t} & \dots & \frac{c_{2N}^2}{|c_1^2|} e^{-(\sigma_1 - \sigma_{2N})t} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ s_k & \frac{c_2^k}{|c_1^k|} e^{-(\sigma_1 - \sigma_2)t} & \frac{c_3^k}{|c_1^k|} e^{-(\sigma_1 - \sigma_3)t} & \dots & \frac{c_{2N}^k}{|c_1^k|} e^{-(\sigma_1 - \sigma_{2N})t} \end{bmatrix} \cdot \begin{bmatrix} \hat{\mathbf{u}}_1 \\ \hat{\mathbf{u}}_2 \\ \vdots \\ \hat{\mathbf{u}}_{2N} \end{bmatrix}$$

with $s_i = \text{sign}(c_1^i)$.

Behavior of GALI_k for chaotic motion

From all determinants appearing in the definition of GALI_k the one that **decreases the slowest** is the one containing the first k columns of the previous matrix:

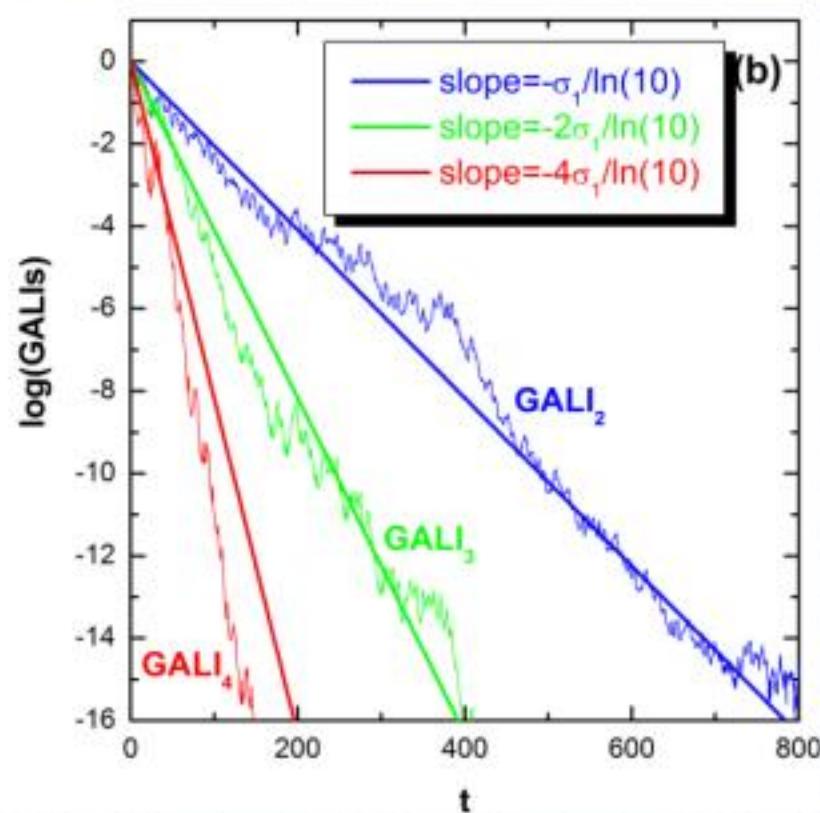
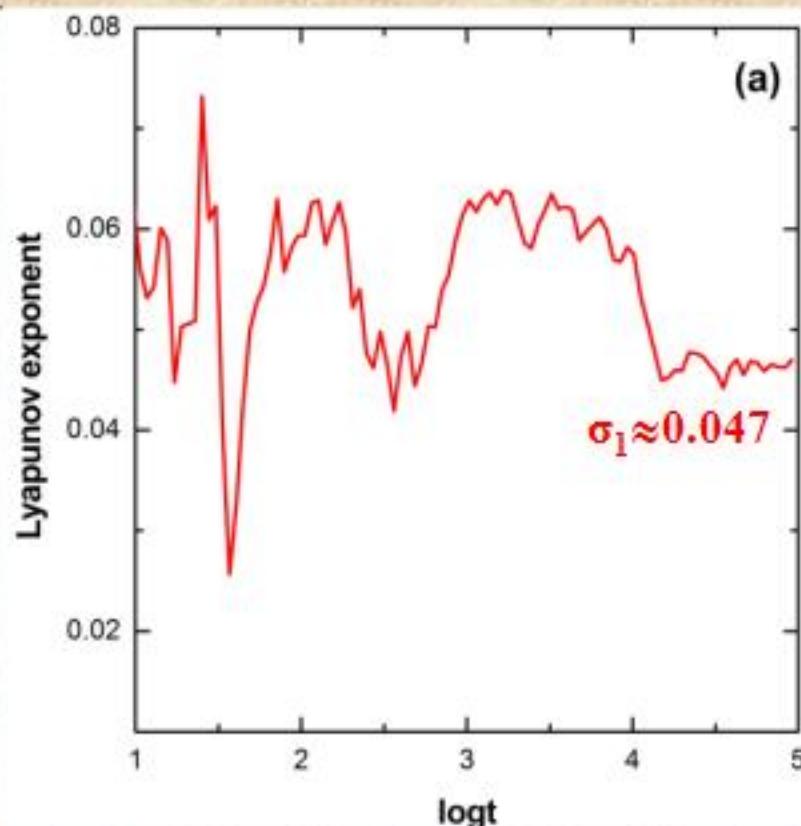
$$\begin{vmatrix} s_1 & \frac{\mathbf{c}_2^1}{|\mathbf{c}_1^1|} e^{-(\sigma_1 - \sigma_2)t} & \frac{\mathbf{c}_3^1}{|\mathbf{c}_1^1|} e^{-(\sigma_1 - \sigma_3)t} & \dots & \frac{\mathbf{c}_k^1}{|\mathbf{c}_1^1|} e^{-(\sigma_1 - \sigma_k)t} \\ s_2 & \frac{\mathbf{c}_2^2}{|\mathbf{c}_1^2|} e^{-(\sigma_1 - \sigma_2)t} & \frac{\mathbf{c}_3^2}{|\mathbf{c}_1^2|} e^{-(\sigma_1 - \sigma_3)t} & \dots & \frac{\mathbf{c}_k^2}{|\mathbf{c}_1^2|} e^{-(\sigma_1 - \sigma_k)t} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ s_k & \frac{\mathbf{c}_2^k}{|\mathbf{c}_1^k|} e^{-(\sigma_1 - \sigma_2)t} & \frac{\mathbf{c}_3^k}{|\mathbf{c}_1^k|} e^{-(\sigma_1 - \sigma_3)t} & \dots & \frac{\mathbf{c}_k^k}{|\mathbf{c}_1^k|} e^{-(\sigma_1 - \sigma_k)t} \end{vmatrix} = \begin{vmatrix} s_1 & \frac{\mathbf{c}_2^1}{|\mathbf{c}_1^1|} & \frac{\mathbf{c}_3^1}{|\mathbf{c}_1^1|} & \dots & \frac{\mathbf{c}_k^1}{|\mathbf{c}_1^1|} \\ s_2 & \frac{\mathbf{c}_2^2}{|\mathbf{c}_1^2|} & \frac{\mathbf{c}_3^2}{|\mathbf{c}_1^2|} & \dots & \frac{\mathbf{c}_k^2}{|\mathbf{c}_1^2|} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ s_k & \frac{\mathbf{c}_2^k}{|\mathbf{c}_1^k|} & \frac{\mathbf{c}_3^k}{|\mathbf{c}_1^k|} & \dots & \frac{\mathbf{c}_k^k}{|\mathbf{c}_1^k|} \end{vmatrix} \cdot e^{-[(\sigma_1 - \sigma_2) + (\sigma_1 - \sigma_3) + \dots + (\sigma_1 - \sigma_k)]t}$$

Thus

$$\text{GALI}_k(t) \propto e^{-[(\sigma_1 - \sigma_2) + (\sigma_1 - \sigma_3) + \dots + (\sigma_1 - \sigma_k)]t}$$

Behavior of GALI_k for chaotic motion

2D Hamiltonian (Hénon-Heiles system)

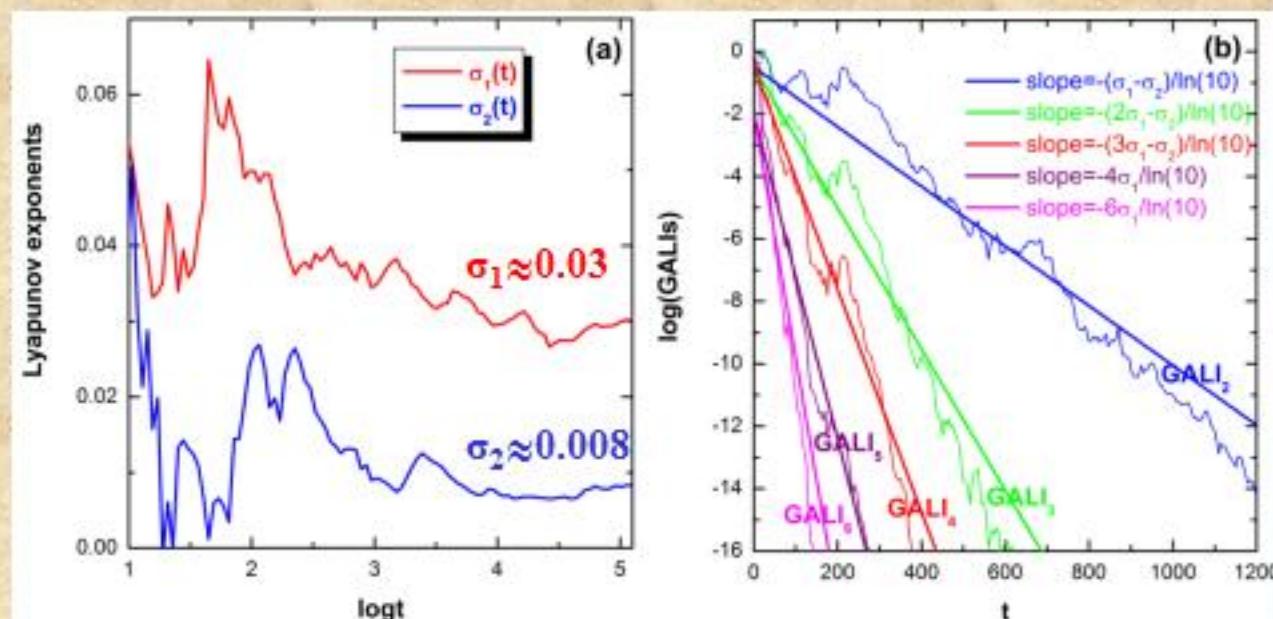


Behavior of GALI_k for chaotic motion

3D system:

$$H_3 = \sum_{i=1}^3 \frac{\omega_i}{2} (q_i^2 + p_i^2) + q_1^2 q_2 + q_1^2 q_3$$

with $\omega_1=1$, $\omega_2=\sqrt{2}$, $\omega_3=\sqrt{3}$, $H_3=0.09$.

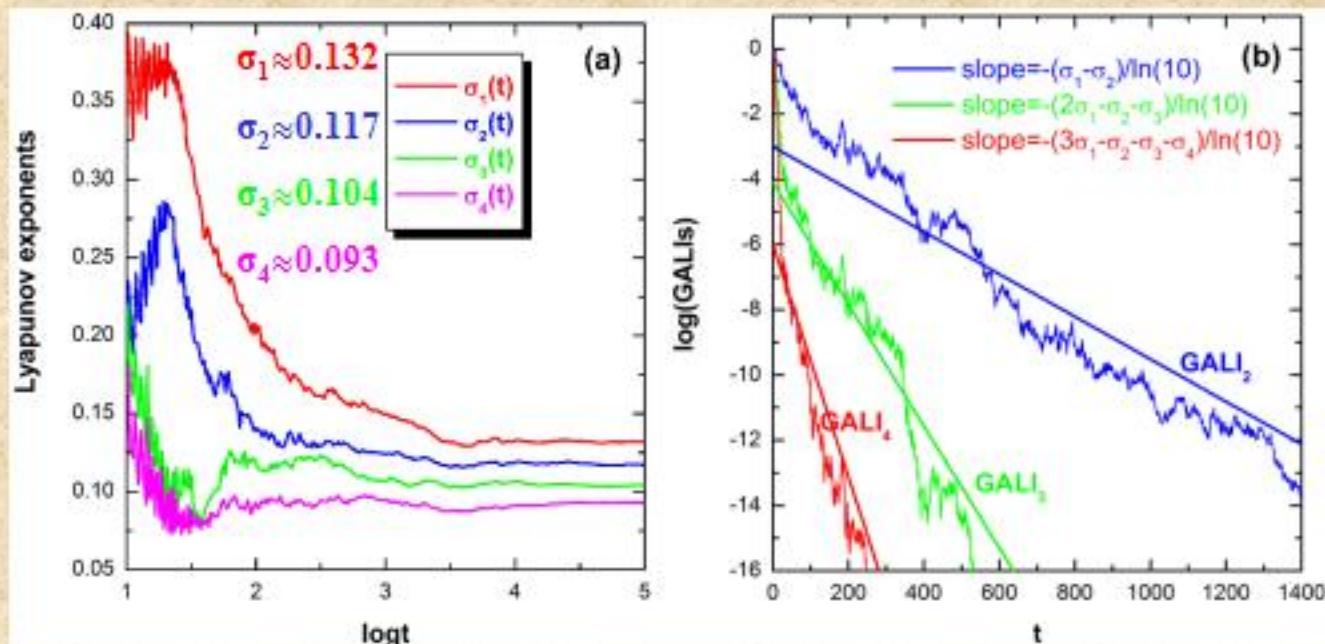


Behavior of GALI_k for chaotic motion

15D Fermi-Pasta-Ulam (FPU) system:

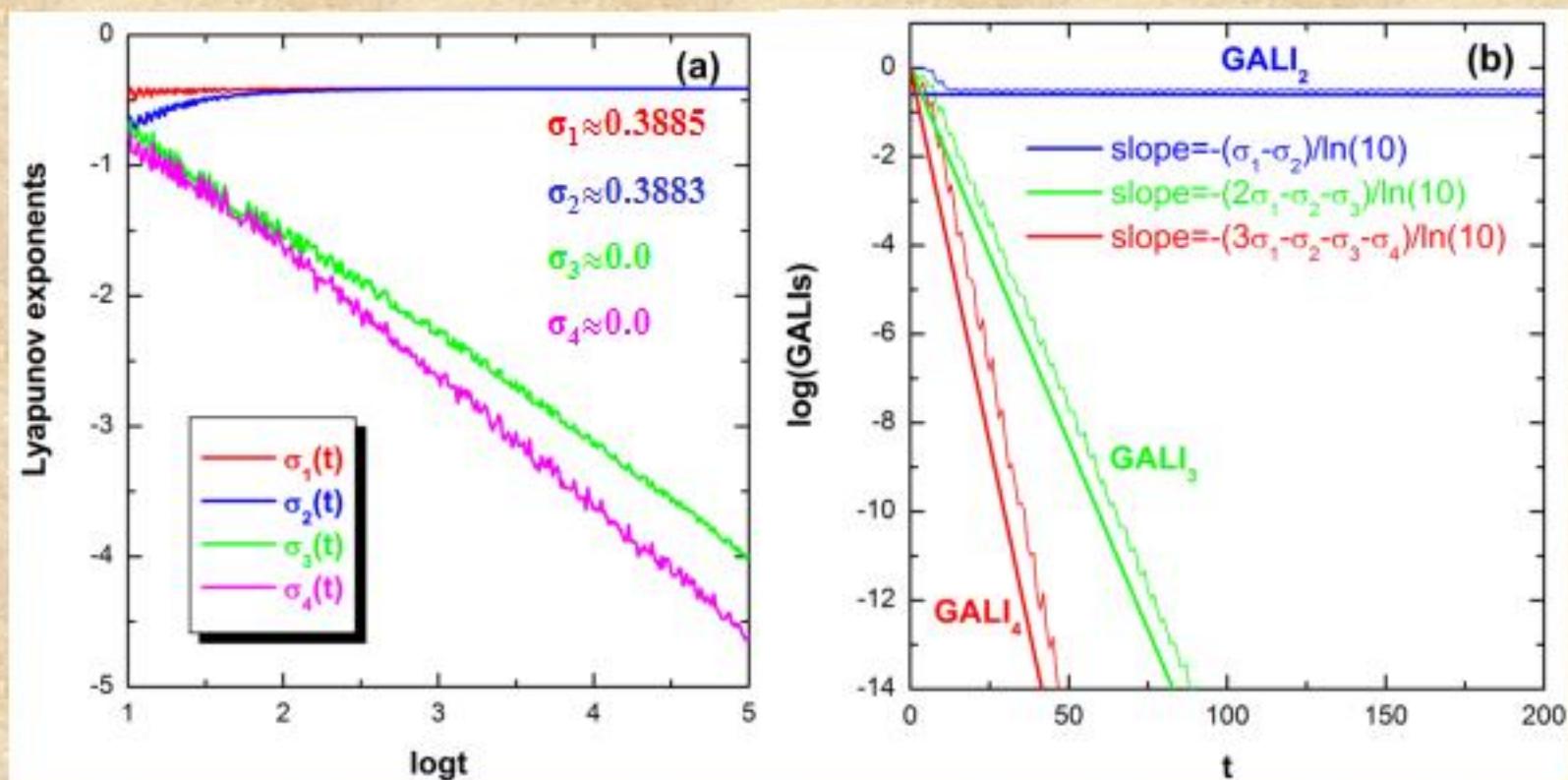
$$H_{15} = \frac{1}{2} \sum_{i=1}^{15} p_i^2 + \sum_{i=1}^{15} \left[\frac{1}{2} (q_{i+1} - q_i)^2 + \frac{1}{4} \beta (q_{i+1} - q_i)^4 \right]$$

with $H_{15}=26.68777$ and $\beta=1.04$.



Behavior of GALI_k for chaotic motion

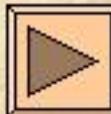
15D Hamiltonian – close to an unstable periodic orbit



Behavior of GALI_k for regular motion

GALI_k remains essentially constant when $2 \leq k \leq N$, while it tends to zero for $N < k \leq 2N$ following a power law, which depends on the number m ($m \leq N$ and $m \leq k$) of deviation vectors that are initially tangent to the torus on which the motion lies:

$$\text{GALI}_k(t) \propto \begin{cases} \text{constant} & \text{if } 2 \leq k \leq N \\ \frac{1}{t^{2(k-N)-m}} & \text{if } N < k \leq 2N \text{ and } 0 \leq m < k - N \\ \frac{1}{t^{k-N}} & \text{if } N < k \leq 2N \text{ and } m \geq k - N \end{cases}$$



Behavior of GALI_k for regular motion

Regular orbits of an N degree of freedom Hamiltonian system lie on **N-dimensional tori**.

Performing a local transformation to action-angle variables we get for the Hamilton's equations of motion:

$$\left. \begin{array}{l} \dot{J}_i = 0 \\ \dot{\theta}_i = \omega_i(J_1, J_2, \dots, J_N) \end{array} \right\} \Rightarrow \begin{array}{l} J_i(t) = J_{i0} \\ \theta_i(t) = \theta_{i0} + \omega_i(J_{10}, J_{20}, \dots, J_{N0}) \cdot t \end{array}, \quad i = 1, 2, \dots, N$$

where J_{i0} , θ_{i0} , $i=1,2,\dots,N$ are the initial conditions.

Behavior of GALI_k for regular motion

The variational equations give:

$$\left. \begin{array}{l} \dot{\xi}_i = 0 \\ \dot{\eta}_i = \sum_{j=1}^N \omega_{ij} \cdot \xi_j \end{array} \right\} \Rightarrow \begin{array}{l} \xi_i(t) = \xi_i(0) \\ \eta_i(t) = \eta_i(0) + \left[\sum_{j=1}^N \omega_{ij} \cdot \xi_j(0) \right] \cdot t \end{array}, \quad i = 1, 2, \dots, N$$

where $\omega_{ij} = \partial \omega_i / \partial J_j|_{J_0}$, $i = 1, 2, \dots, N$ are constants.

Using as a basis of the $2N$ -dimensional tangent space of the flow the $2N$ unit vectors $\{\hat{u}_1, \hat{u}_2, \dots, \hat{u}_{2N}\}$ such that the first N of them, $\hat{u}_1, \hat{u}_2, \dots, \hat{u}_N$ correspond to the N action variables and the N remaining ones,

$\hat{u}_{N+1}, \hat{u}_{N+2}, \dots, \hat{u}_{2N}$ to the conjugate angle variables, we write any unit deviation vector as:

$$\hat{v}_i(t) = \frac{1}{\|\hat{v}_i(t)\|} \left[\sum_{j=1}^N \xi_j^i(0) \hat{u}_j + \sum_{j=1}^N \left(\eta_j^i(0) + \sum_{k=1}^N \omega_{kj} \xi_k^i(0) t \right) \hat{u}_{N+j} \right]$$

with $\|\hat{v}_i(t)\| \propto t$

Behavior of GALI_k for regular motion

For k deviation vectors we have:

$$\begin{bmatrix} \hat{\mathbf{v}}_1 \\ \hat{\mathbf{v}}_2 \\ \vdots \\ \hat{\mathbf{v}}_k \end{bmatrix} = \frac{\mathbf{1}}{\prod_{m=1}^k \|\mathbf{v}_m(t)\|}.$$

$$\begin{bmatrix} \xi_1^1(0) & \dots & \xi_N^1(0) & \eta_1^1(0) + \sum_{m=1}^N \omega_{1m} \xi_m^1(0) t & \dots & \eta_N^1(0) + \sum_{m=1}^N \omega_{Nm} \xi_m^1(0) t \\ \xi_1^2(0) & \dots & \xi_N^2(0) & \eta_1^2(0) + \sum_{m=1}^N \omega_{1m} \xi_m^2(0) t & \dots & \eta_N^2(0) + \sum_{m=1}^N \omega_{Nm} \xi_m^2(0) t \\ \vdots & & \vdots & \vdots & & \vdots \\ \xi_1^k(0) & \dots & \xi_N^k(0) & \eta_1^k(0) + \sum_{m=1}^N \omega_{1m} \xi_m^k(0) t & \dots & \eta_N^k(0) + \sum_{m=1}^N \omega_{Nm} \xi_m^k(0) t \end{bmatrix} \cdot \begin{bmatrix} \hat{\mathbf{u}}_1 \\ \hat{\mathbf{u}}_2 \\ \vdots \\ \hat{\mathbf{u}}_{2N} \end{bmatrix}$$

Behavior of GALI_k for regular motion

For $2 \leq k \leq N$ the slowest decreasing determinants are the ones whose k columns are chosen among the last N columns of the evolution matrix

$$\frac{1}{\prod_{m=1}^k \|v_m(t)\|} \begin{vmatrix} \eta_{j_1}^1(0) + \sum_{m=1}^N \omega_{jm} \xi_m^1(0) t & \dots & \eta_{j_k}^1(0) + \sum_{m=1}^N \omega_{jm} \xi_m^1(0) t \\ \eta_{j_1}^2(0) + \sum_{m=1}^N \omega_{jm} \xi_m^2(0) t & \dots & \eta_{j_k}^2(0) + \sum_{m=1}^N \omega_{jm} \xi_m^2(0) t \\ \vdots & & \vdots \\ \eta_{j_1}^k(0) + \sum_{m=1}^N \omega_{jm} \xi_m^k(0) t & \dots & \eta_{j_k}^k(0) + \sum_{m=1}^N \omega_{jm} \xi_m^k(0) t \end{vmatrix} \propto$$

$$\propto \frac{1}{t^k} \begin{vmatrix} \omega_{j_1 m_1} \xi_{m_1}^1(0) t & \dots & \omega_{j_k m_1} \xi_{m_1}^1(0) t \\ \omega_{j_1 m_1} \xi_{m_1}^2(0) t & \dots & \omega_{j_k m_1} \xi_{m_1}^2(0) t \\ \vdots & & \vdots \\ \omega_{j_1 m_1} \xi_{m_1}^k(0) t & \dots & \omega_{j_k m_1} \xi_{m_1}^k(0) t \end{vmatrix} \propto \frac{1}{t^k} \cdot t^k \begin{vmatrix} \omega_{j_1 m_1} \xi_{m_1}^1(0) & \dots & \omega_{j_k m_1} \xi_{m_1}^1(0) \\ \omega_{j_1 m_1} \xi_{m_1}^2(0) & \dots & \omega_{j_k m_1} \xi_{m_1}^2(0) \\ \vdots & & \vdots \\ \omega_{j_1 m_1} \xi_{m_1}^k(0) & \dots & \omega_{j_k m_1} \xi_{m_1}^k(0) \end{vmatrix} \approx \text{constant}$$

Behavior of GALI_k for regular motion

For $N < k \leq 2N$ the slowest decreasing determinants are the ones containing the last N columns of the evolution matrix

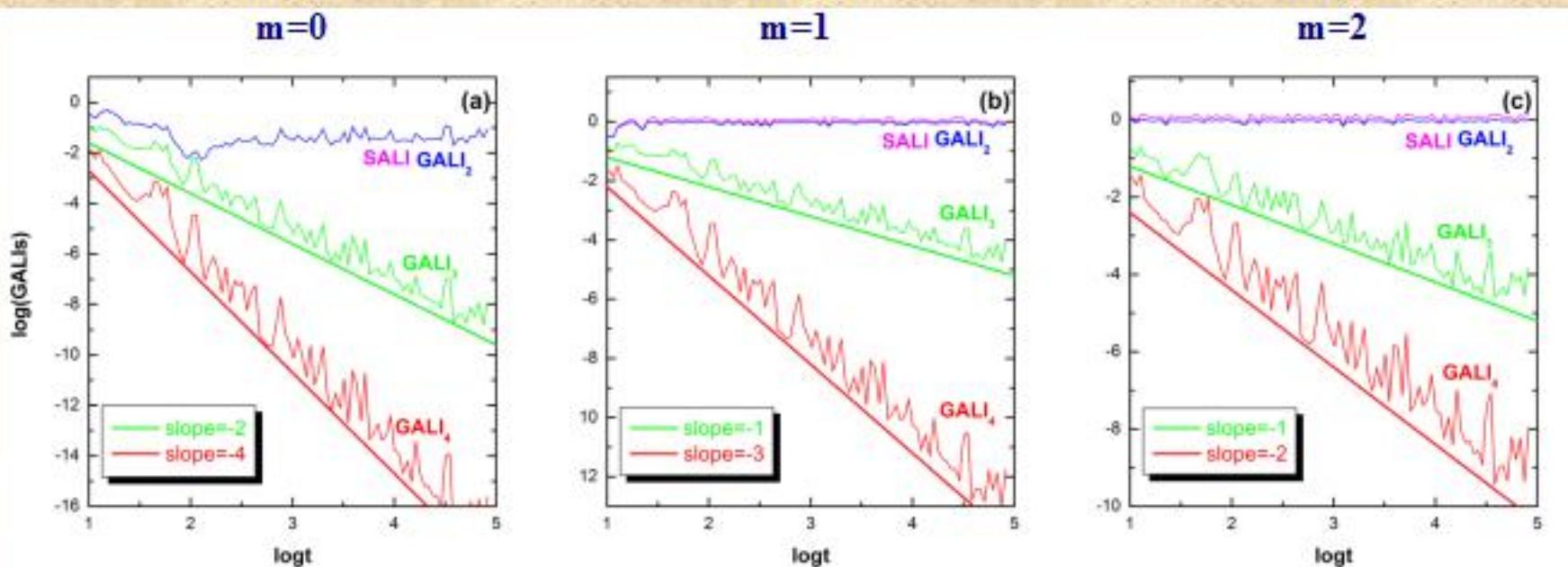
$$\frac{1}{\prod_{m=1}^k \|v_m(t)\|} \begin{vmatrix} \xi_{i_1}^1(0) & \dots & \xi_{i_{k-N}}^1(0) & \eta_i^1(0) + \sum_{m=1}^N \omega_{lm} \xi_m^1(0) t & \dots & \eta_N^1(0) + \sum_{m=1}^N \omega_{Nm} \xi_m^1(0) t \\ \xi_{i_1}^2(0) & \dots & \xi_{i_{k-N}}^2(0) & \eta_i^2(0) + \sum_{m=1}^N \omega_{lm} \xi_m^2(0) t & \dots & \eta_N^2(0) + \sum_{m=1}^N \omega_{Nm} \xi_m^2(0) t \\ \vdots & & \vdots & \vdots & & \vdots \\ \xi_{i_1}^k(0) & \dots & \xi_{i_{k-N}}^k(0) & \eta_i^k(0) + \sum_{m=1}^N \omega_{lm} \xi_m^k(0) t & \dots & \eta_N^k(0) + \sum_{m=1}^N \omega_{Nm} \xi_m^k(0) t \end{vmatrix} \propto$$

$$\propto \frac{1}{t^k} \begin{vmatrix} \underbrace{\xi_{i_1}^1(0) & \dots & \xi_{i_{k-N}}^1(0)}_{k-N \text{ columns}} & \underbrace{\eta_i^1(0) & \dots & \eta_{i_{k-N}}^1(0)}_{k-N \text{ columns}} & \underbrace{\omega_{i_{k-N+1} m_1} \xi_{m_1}^1(0) t & \dots & \omega_{i_N m_{2N}} \xi_{m_1}^1(0) t}_{2N-k \text{ columns}} \\ \xi_{i_1}^2(0) & \dots & \xi_{i_{k-N}}^2(0) & \eta_i^2(0) & \dots & \eta_{i_{k-N}}^2(0) & \omega_{i_{k-N+1} m_1} \xi_{m_1}^2(0) t & \dots & \omega_{i_N m_{2N}} \xi_{m_1}^2(0) t \\ \vdots & & \vdots & \vdots & & \vdots & \vdots & & \vdots \\ \xi_{i_1}^k(0) & \dots & \xi_{i_{k-N}}^k(0) & \eta_i^k(0) & \dots & \eta_{i_{k-N}}^k(0) & \omega_{i_{k-N+1} m_1} \xi_{m_1}^k(0) t & \dots & \omega_{i_N m_{2N}} \xi_{m_1}^k(0) t \end{vmatrix} \propto \frac{1}{t^k} \cdot t^{2N-k} \approx \frac{1}{t^{2(k-N)}}$$

Behavior of GALI_k for regular motion

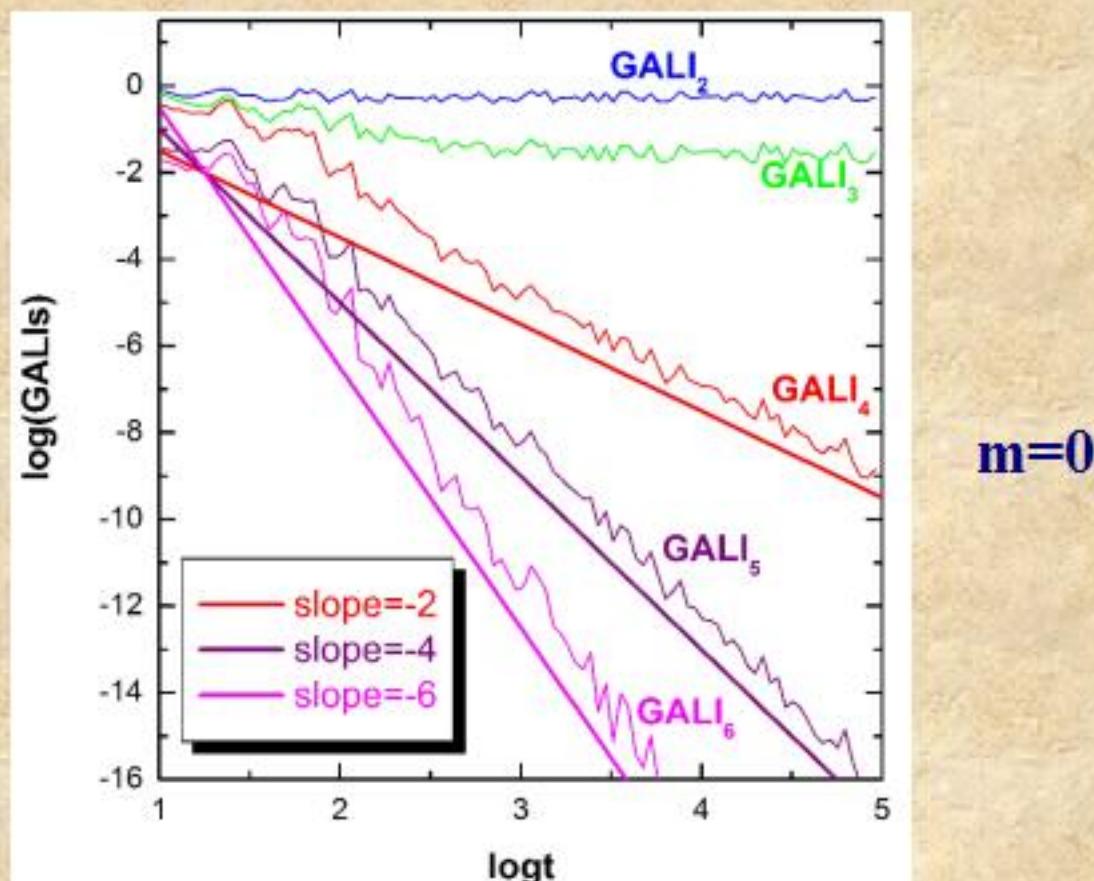
2D Hamiltonian
(Hénon-Heiles system) $N=2$

$$\text{GALI}_k(t) \propto \begin{cases} \text{constant} & \text{if } 2 \leq k \leq N \\ \frac{1}{t^{2(k-N)-m}} & \text{if } N < k \leq 2N \text{ and } 0 \leq m < k-N \\ \frac{1}{t^{k-N}} & \text{if } N < k \leq 2N \text{ and } m \geq k-N \end{cases}$$



Behavior of GALI_k for regular motion

3D Hamiltonian



$m=0$

Behavior of GALI_k

Chaotic motion:

SALI → 0 exponential decay

$$\text{GALI}_k(t) \propto e^{-[(\sigma_1 - \sigma_2) + (\sigma_1 - \sigma_3) + \dots + (\sigma_1 - \sigma_k)]t}$$

Regular motion:

SALI → constant ≠ 0 or SALI → 0 power law decay

$$\text{GALI}_k(t) \propto \begin{cases} \text{constant} & \text{if } 2 \leq k \leq N \\ \frac{1}{t^{2(k-N)-m}} & \text{if } N < k \leq 2N \text{ and } 0 \leq m < k - N \\ \frac{1}{t^{k-N}} & \text{if } N < k \leq 2N \text{ and } m \geq k - N \end{cases}$$

Global dynamics

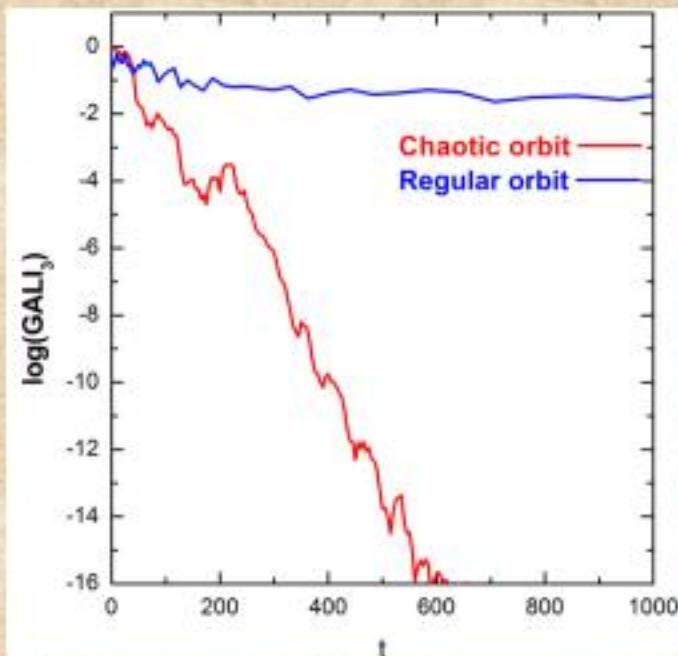
- GALI_2 (practically equivalent to the use of SALI)

- GALI_N

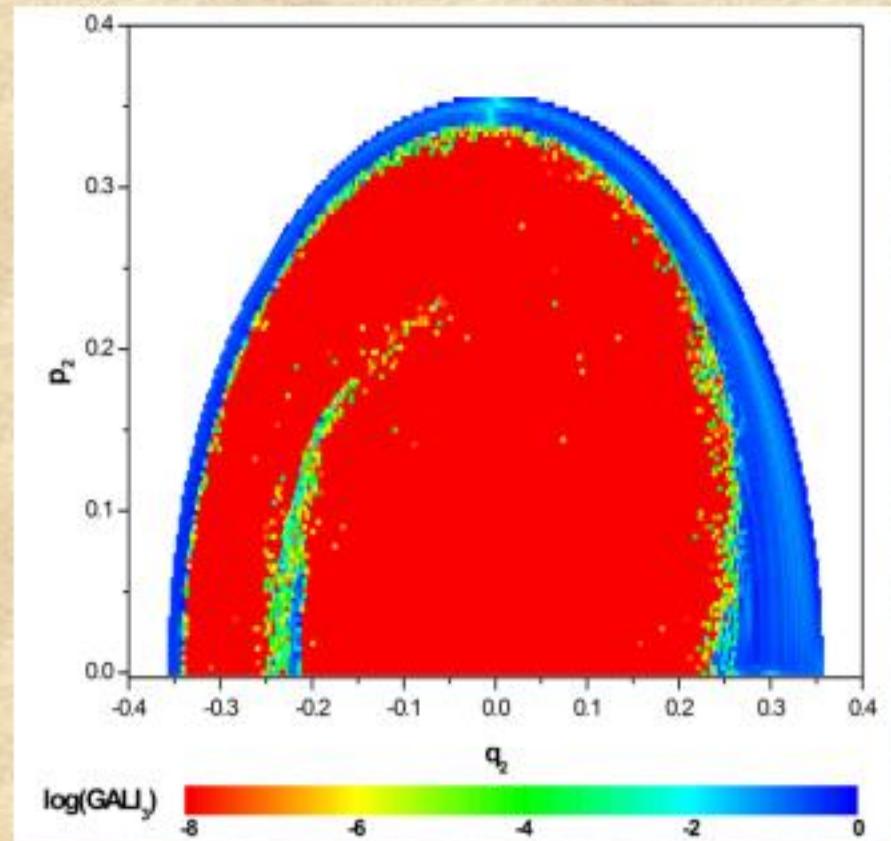
Chaotic motion: $\text{GALI}_N \rightarrow 0$
(exponential decay)

Regular motion:

$\text{GALI}_N \rightarrow \text{constant} \neq 0$



3D Hamiltonian
Subspace $q_3=p_3=0$, $p_2 \geq 0$ for $t=1000$.



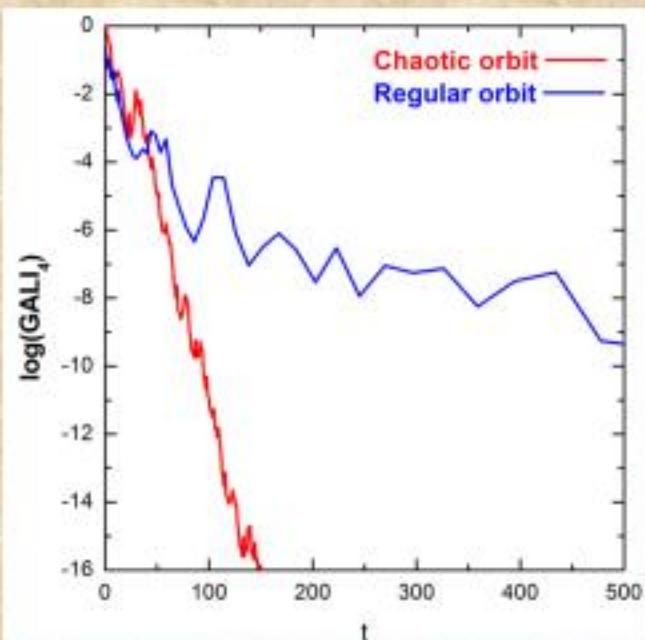
Global dynamics

GALI_k with k>N

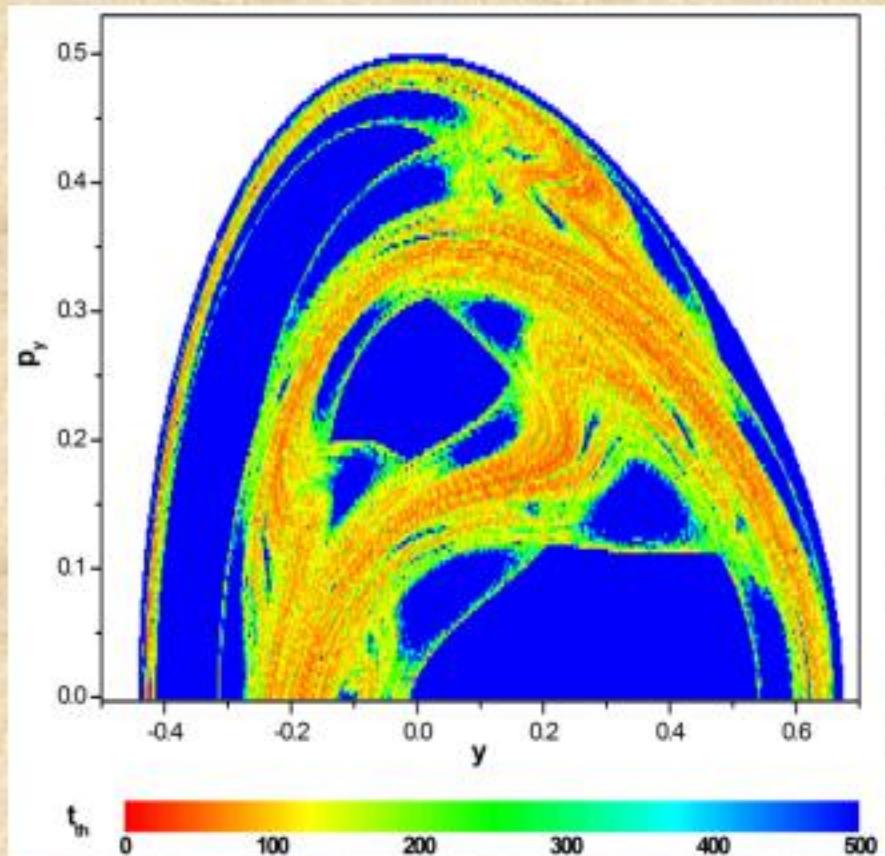
The index tends to zero both for regular and chaotic orbits but with completely different time rates:

Chaotic motion: exponential decay

Regular motion: power law



2D Hamiltonian (Hénon-Heiles)
Time needed for $\text{GALI}_4 < 10^{-12}$



References - 1

• Hamiltonian systems and symplectic maps

- ✓ Lieberman A. J. & Lichtenberg M. A., 1992, Regular and Chaotic Dynamics, Springer.
- ✓ Cvitanović P., Artuso R., Dahlqvist P., Mainieri R., Tanner G., Vattay G., Whelan N. & Wirsba A., 2004, Chaos – Classical and Quantum, <http://chaosbook.org/>

• Lyapunov exponents

- ✓ Oseledec V I, 1968 *Trans. Moscow Math. Soc.* **19** 197
- ✓ Benettin G, Galgani L and Strelcyn J-M 1976 *Phys. Rev. A* **14** 2338
- ✓ Pesin Y B, 1977 *Russian Math. Surveys* **32** 55
- ✓ Benettin G, Galgani L, Giorgilli A and Strelcyn J-M 1980 *Meccanica* March **9**
- ✓ Benettin G, Galgani L, Giorgilli A and Strelcyn J-M 1980 *Meccanica* March **21**
- ✓ Greene J M and Kim J-S *Physica D* **24** 213
- ✓ Wolf A, Swift J B, Swinney H L and Vastano J A 1985 *Physica D* **16** 285
- ✓ Bridges T J and Reich S *Physica D* **156** 219

• Frequency Analysis

- ✓ Laskar J 1988 *Astron. Astrophys.* **198** 341
- ✓ Laskar J 1990 *Icarus* **88** 266
- ✓ Laskar J, Froeschlé C and Celletti A 1992 *Physica D* **56** 253
- ✓ Laskar J 1993 *Physica D* **67** 257
- ✓ Dumas H S and Laskar J 1993 *Phys. Rev. Lett.* **70** 2975
- ✓ Papaphilippou Y and Laskar J 1996 *Astron. Astrophys.* **307** 427
- ✓ Papaphilippou Y and Laskar J 1998 *Astron. Astrophys.* **329** 451
- ✓ Laskar J 1999 *Hamiltonian systems with three or more degrees of freedom* (ed. Simo C / Plenum Press) p 134
- ✓ Robutel P and Laskar J 2001 *Icarus* **152** 4
- ✓ Robutel P and Gabern F 2006 *MNRAS* **372** 1463

References - 2

• SALI - GALI

- ✓ Skokos Ch 2001 *J. Phys. A* **34** 10029
- ✓ Skokos Ch, Antonopoulos Ch, Bountis T C and Vrahatis M N 2003 *Prog. Theor. Phys. Suppl.* **150** 439
- ✓ Skokos Ch, Antonopoulos Ch, Bountis T C and Vrahatis MN 2004 *J. Phys. A* **37** 6269
- ✓ Skokos Ch, Bountis T C and Antonopoulos Ch 2005 (submitted)
- ✓ Szell A, Erdi B, Sando Zs and Steves B 2004 *Mon. Not. R. Astron. Soc.* **347** 380
- ✓ Kalapotharakos C Voglis N and Contopoulos G 2004 *Astron. Astroph.*, **428** 905
- ✓ Manos T and Athanassoula E 2005 in ‘SF2A-2005: Semaine de l’Astrophysique Française’, eds. Casoli F Contini T Hameury J.M and Pagani L, EDP-Sciences Conference Series, 631
- ✓ Bountis T C and Skokos Ch 2006 *Nucl. Inst. Meth. Phys. Res. A* **561** 173
- ✓ de Assis L P G Helayél-Neto J A Haas F and Nogueira A L M A 2006 Fifth International Conference on Mathematical Methods in Physics – IC2006, Proceedings of Science, PoS(IC2006)059
- ✓ Capuzzo-Dolcetta R, Leccese L, Merritt D and Vicari A 2006 astro-ph/0611205

• FLI

- ✓ Froeschlé C, Lega E and Gonczi R 1997 *Celest. Mech. Dyn. Astron.* **67** 41
- ✓ Froeschlé C, Gonczi R and Lega E 1997 *Planet. Space Sci.* **45** 881
- ✓ Fouchard M, Lega E, Froeschlé Ch and Froeschlé C 2002 *Celest. Mech. Dyn. Astron.* **83** 205
- ✓ Guzzo M, Lega E and Froeschlé C 2002 *Physica D* **163** 1
- ✓ Barrio R 2005 *Chaos Sol. Fract.* **25** 71

References - 3

•MEGNO

- ✓ Cincotta P M and Simo 2000 *Astron. Astroph. Suppl. Ser.* **147** 205
- ✓ Cincotta P M, Giordano C M and Simo C 2003 *Physica D* **182** 151

•RLI

- ✓ Sandor Zs Erdi B Szell A and Funk B 2004 *Celest. Mech. Dyn. Astron.* **90** 127
- ✓ Erdi B and Sandor Zs 2005 *Celest. Mech. Dyn. Astron.* **92** 113

•Dynamical spectra

- ✓ Froeschle C, Froeschle Ch and Lohinger E 1993 *Celest. Mech. Dyn. Astron.* **56** 307
- ✓ Lohinger E, Froeschle C and Dvorak R 1993 *Celest. Mech. Dyn. Astron.* **56** 315
- ✓ Voglis N and Contopoulos G 1994 *J. Phys. A* **27** 4899
- ✓ Contopoulos G and Voglis N 1996 *Celest. Mech. Dyn. Astron.* **64** 1
- ✓ Froeschle C and Lega E 1998 *Astron. Astroph.* **334** 355
- ✓ Voglis N Contopoulos G and Efthymiopoulos C 1999 *Celest. Mech. Dyn. Astron.* **73** 211
- ✓ Vozikis Ch L, Varvoglis H and Tsiganis K 2000 *Astron. Astroph.* **359** 386

•Other methods

- ✓ Voyatzis G and Ichtiaoglou S 1992 *J. Phys. A* **25** 5931
- ✓ Kotoulas T and Voyatzis G 2004 *Celest. Mech. Dyn. Astron.* **88** 343
- ✓ Gottwald G A and Melbourne I 2004 *Proc. Roy. Soc. London A* **460** 603
- ✓ Howard J E 2005 *Celest. Mech. Dyn. Astron.* **92** 219
- ✓ Sideris I V 2006 *Phys. Rev. E* **73** 066217